

# Recent Progress from the DEAP-3600 Dark Matter Direct Detection Experiment

Jocelyn Monroe,  
Royal Holloway University of London

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## Outline

1. DEAP-3600 Detector
2. Experimental Technique
3. Status and Outlook



# DEAP-3600 Collaboration

**University of Alberta** D. Grant, P. Gorel, A. Hallin, J. Tang, J. Soukup, C. Ng, B. Beltran, J. Bueno, K. Olsen, R. Chouinard, T. McElroy, S. Crothers, S. Liu, P. Davis, and A. Viangreiro

**Carleton University** K. Graham, C. Ouellet, K. Brown

**Queen's University** M. Boulay, B. Cai, J. Bonatt, B. Broerman, D. Bearse, K. Dering, M. Chen, S. Florian, R. Gagnon, P. Giampa, V.V. Golovko, P. Harvey, M. Kuzniak, J.J. Lidgard, A. McDonald, C. Nantais, A.J. Noble, E. O'Dwyer, P. Pasuthip, L. Veloce, W. Rau, T. Sonley, P. Skensved, M. Ward

**SNOLAB/Laurentian** B. Cleveland, F. Duncan, R. Ford, C.J. Jillings, T. Pollmann

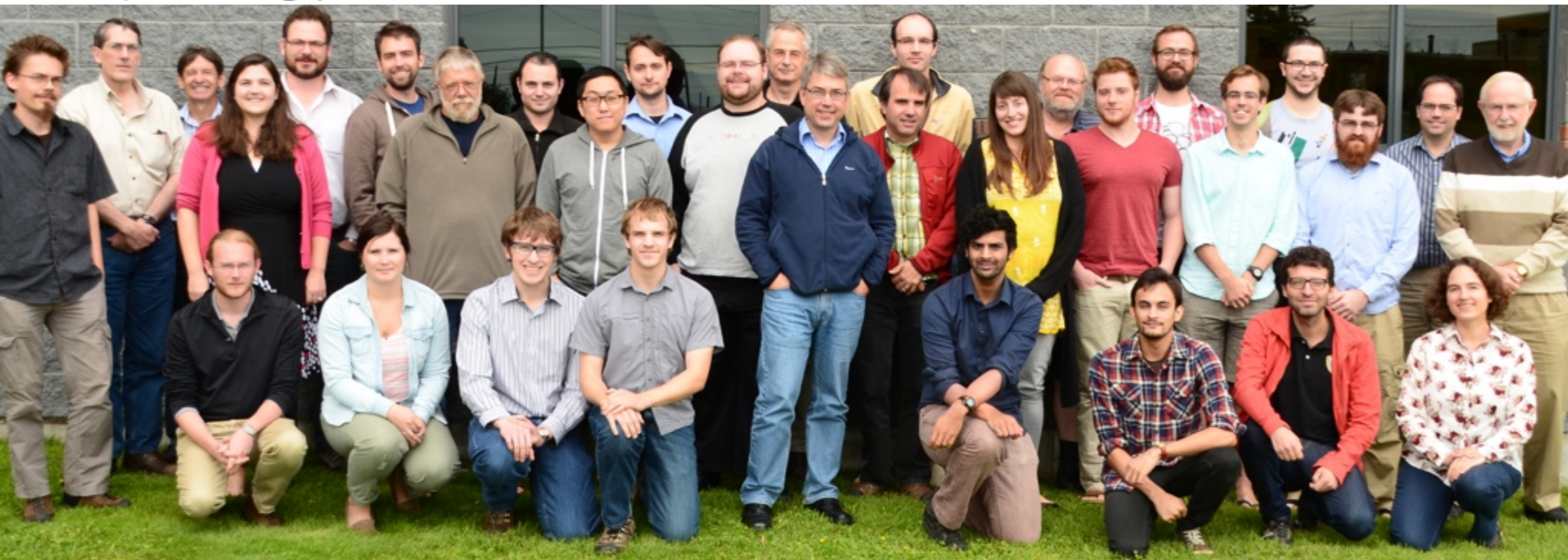
**SNOLAB** I. Lawson, K. McFarlane, P. Liimatainen, O. Li, E. Vazquez Jauregui

**TRIUMF** F. Retiere, P-A. Amaudruz, D. Bishop, S. Chan, C. Lim, A. Muir, C. Ohlmann, K. Olchanski, V. Strickland

**Rutherford Appleton Laboratory** P. Majewski

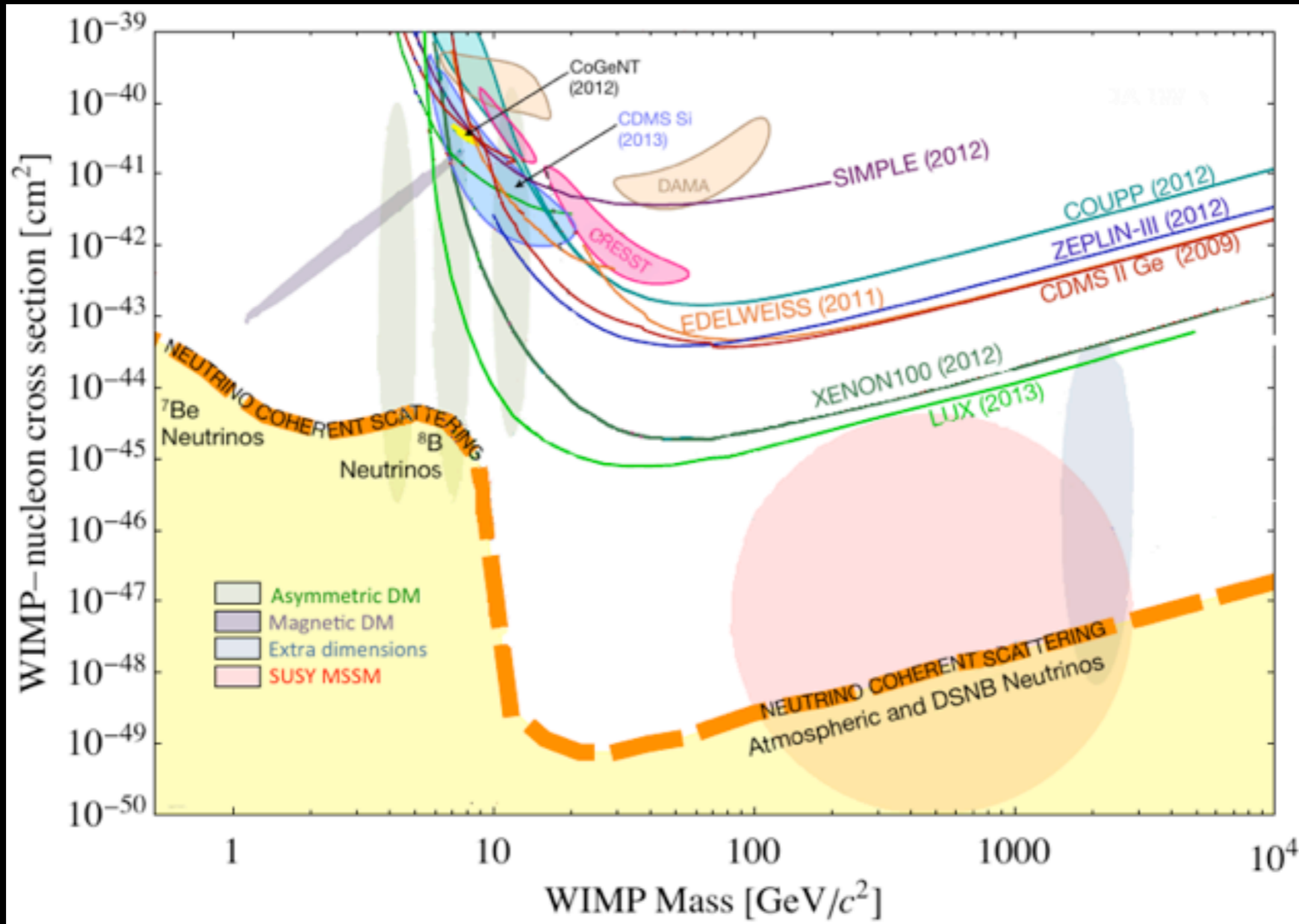
**Royal Holloway University of London** J. Monroe, J. Walding, A. Butcher, N. Seeburn

**University of Sussex** S.J.M. Peeters





# The Low-Background Frontier: Status of SI Searches



← 1 event/  
kg/day

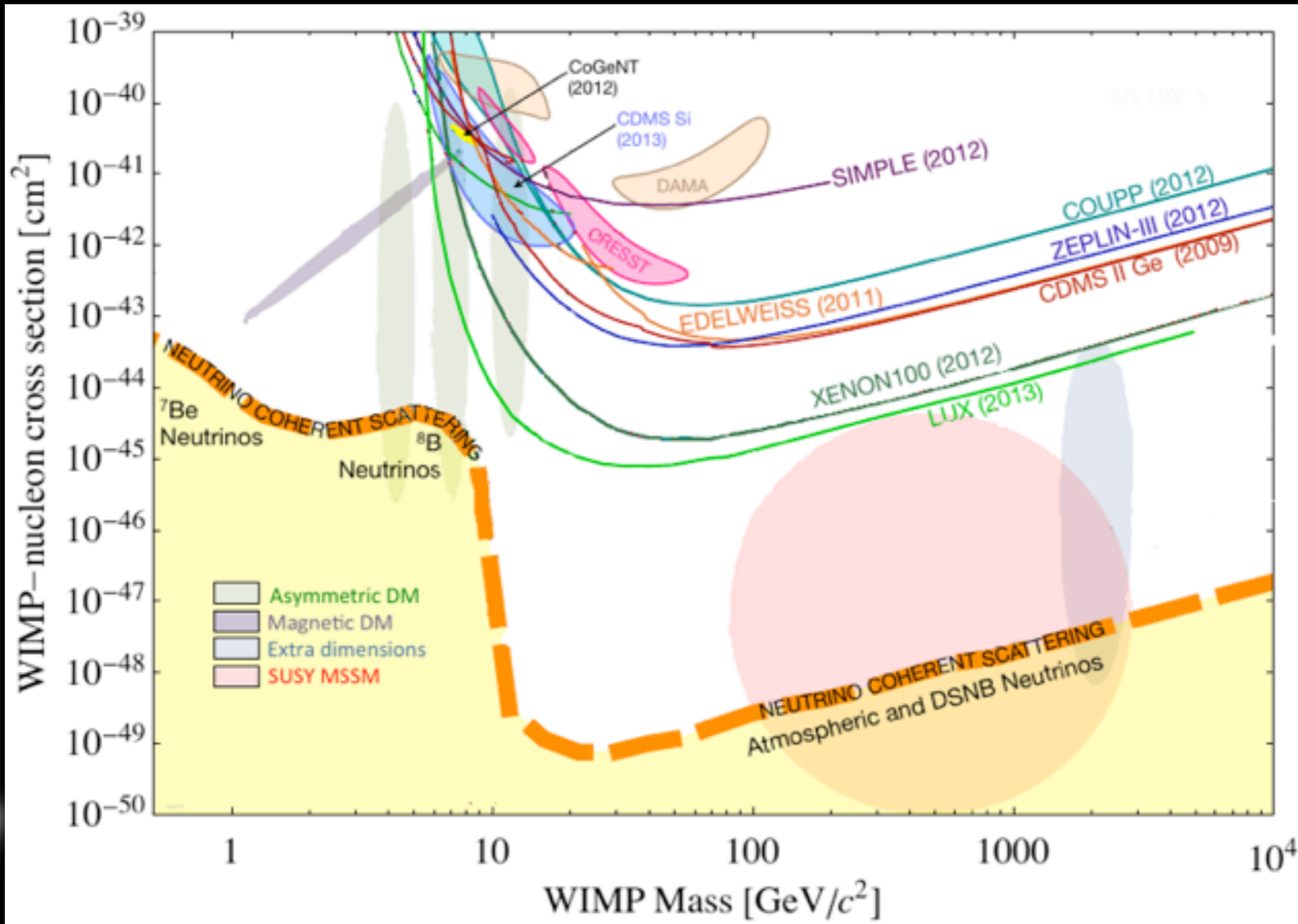
← 1 event/  
100kg/day

← 1 event/  
100 kg/  
100 days



# The Low-Background Frontier: Status of SI Searches

Scalability of Detector Technology



←  $1 \text{ event/kg/day}$

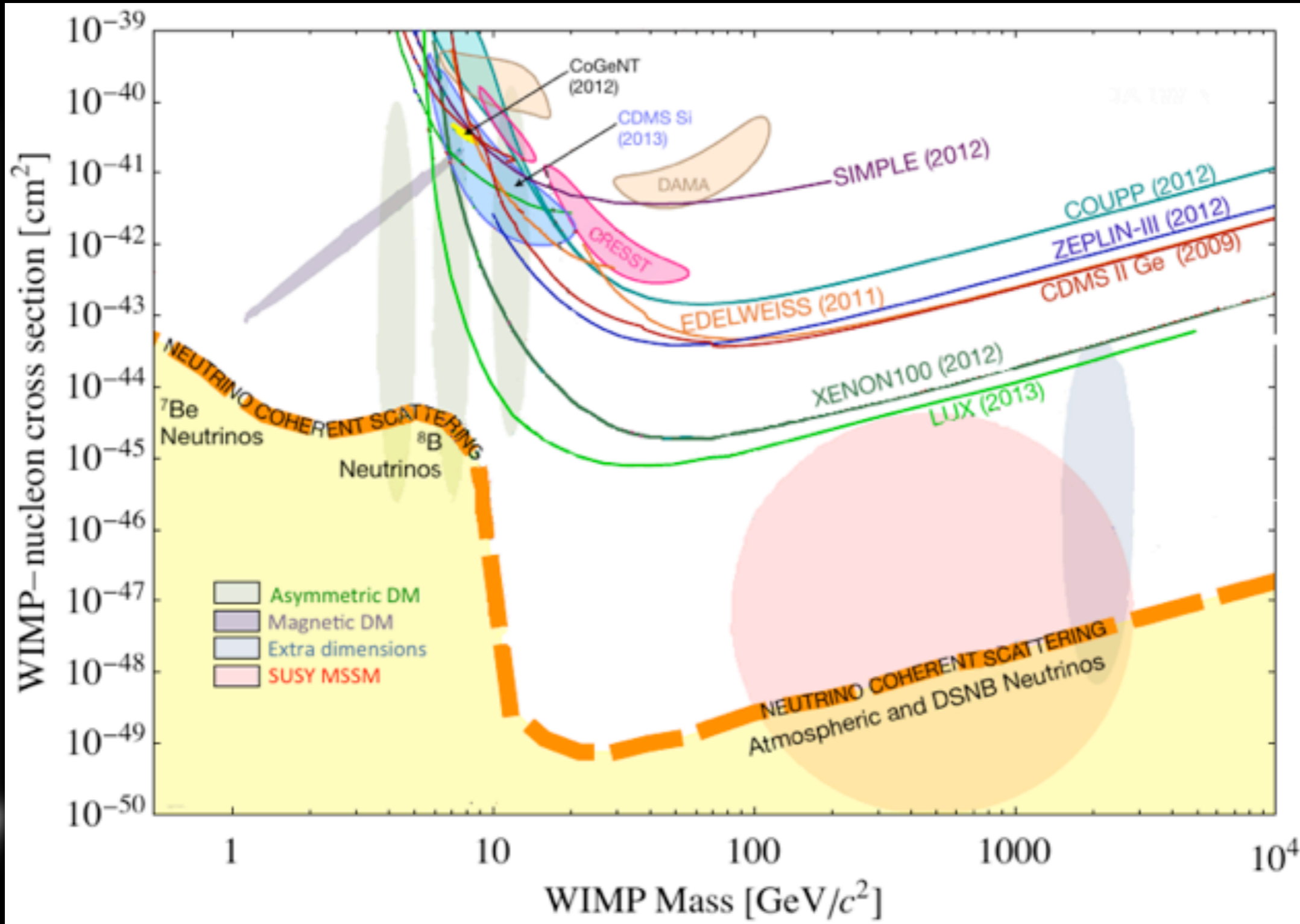
←  $1 \text{ event/100kg/day}$

←  $1 \text{ event/100 kg/100 days}$





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Scalability of Detector Technology

New Techniques for Backgrounds

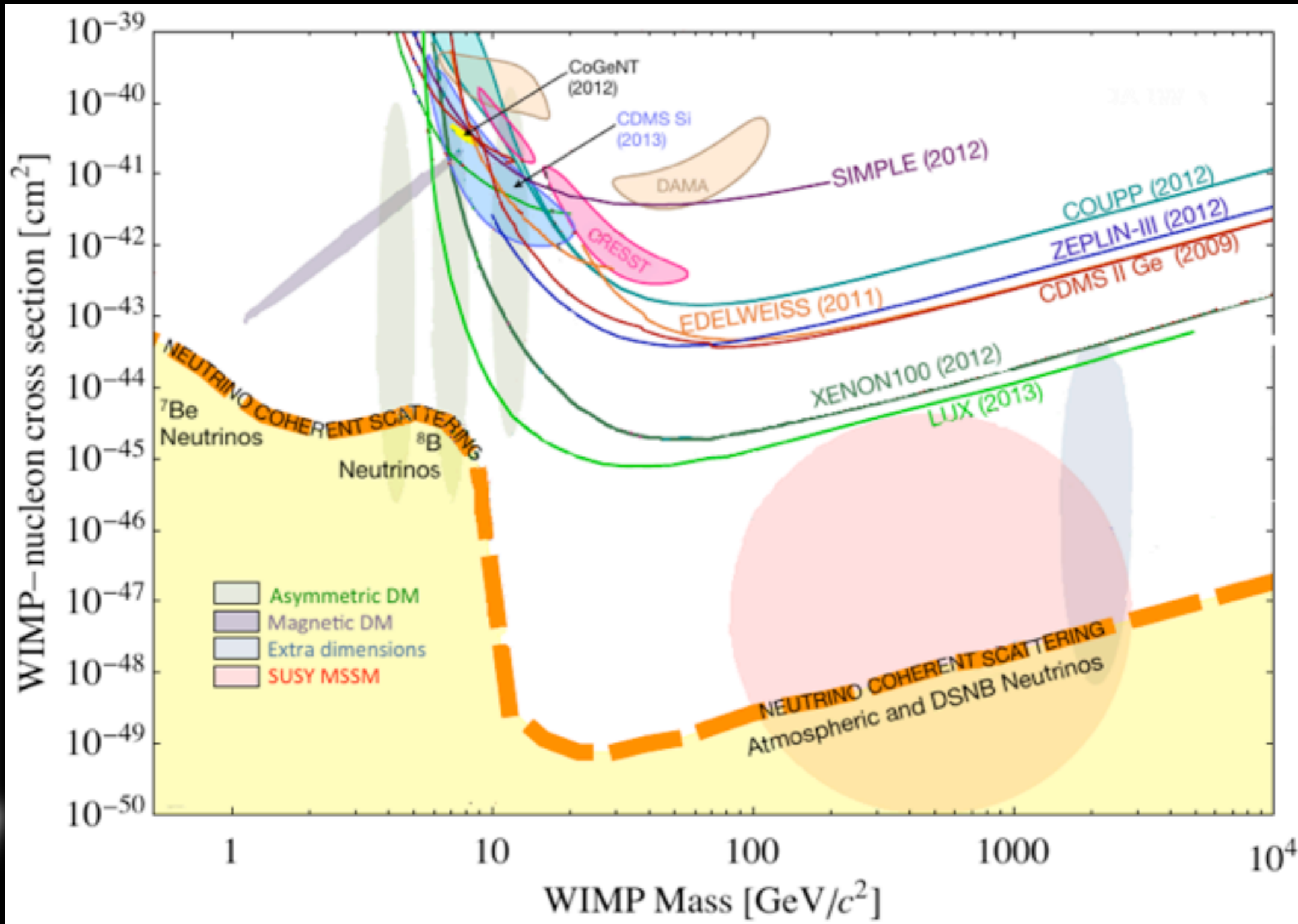
1 event/  
kg/day

1 event/  
100kg/day

1 event/  
100 kg/  
100 days



# The Low-Background Frontier: Status of SI Searches



Scalability of Detector Technology

New Techniques for Backgrounds

1 event/  
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100kg/day

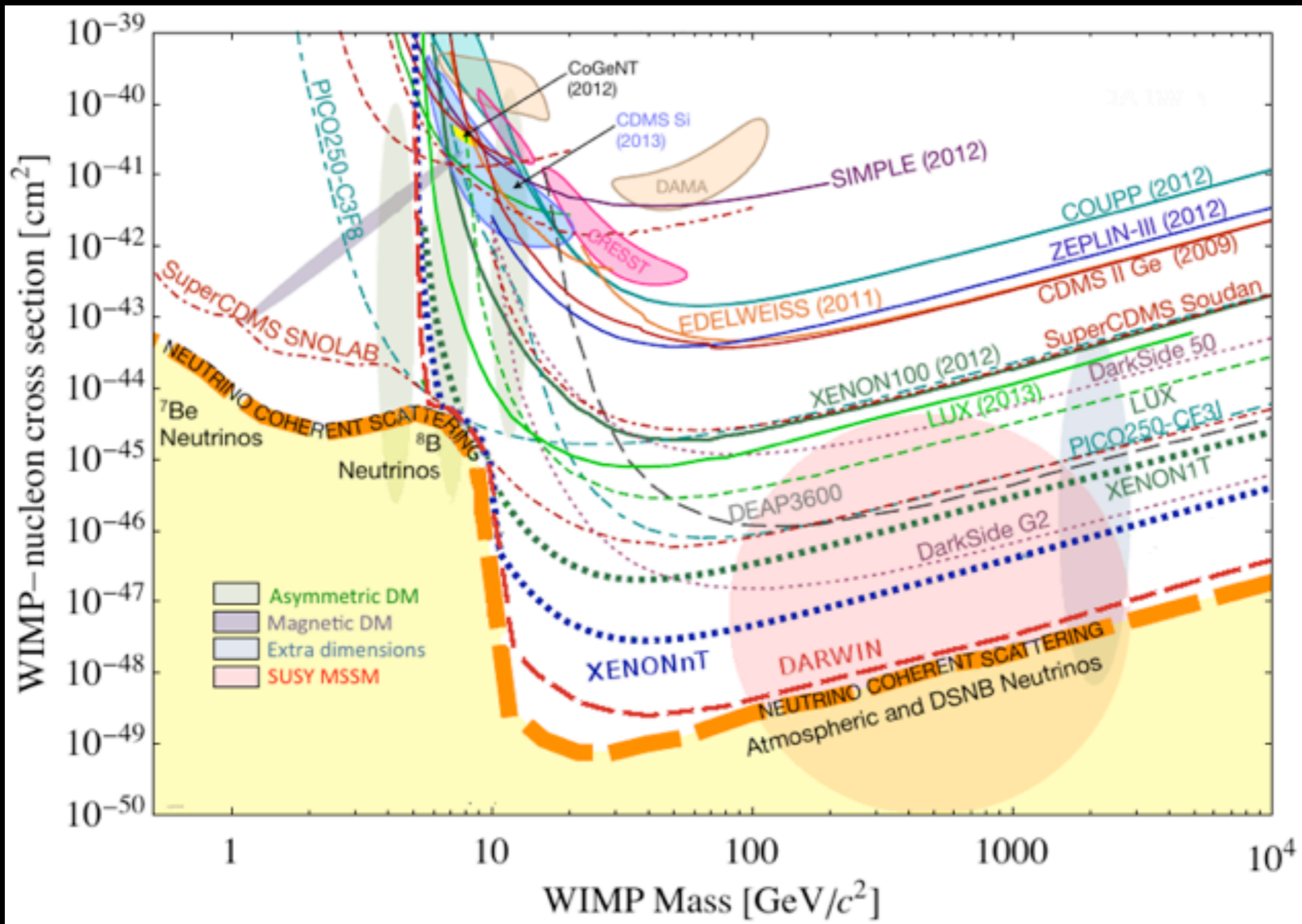
1 event/  
100 kg/  
100 days

Complementary with High-Energy Frontier





# The Low-Background Frontier: Prospects



under  
construction  
proposed

so far:  $<1$  event at  $1\text{E-}45\text{ cm}^2$ , therefore need minimum  $1\text{E-}47\text{ cm}^2$  sensitivity for 100 events to measure  $M_\chi, \sigma$



Neutrino Lesson:

key to scalability is  
large, open volume  
simple detector design

DEAP/CLEAN Strategy:

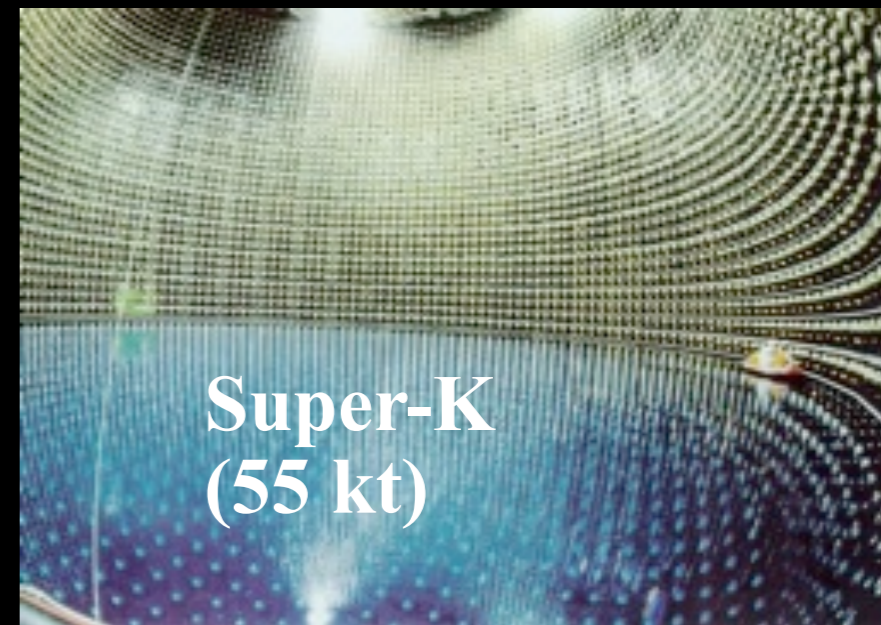
draw on design successes of  
large neutrino experiments

cross section (cm<sup>2</sup>)

detector mass (ktonnes)

10<sup>-45</sup>

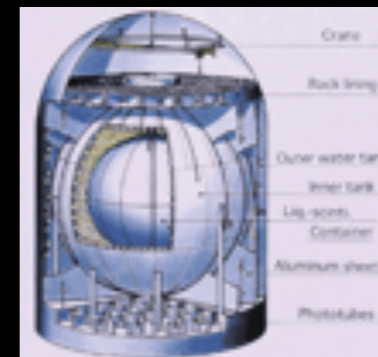
100



30

10<sup>-44</sup>

10

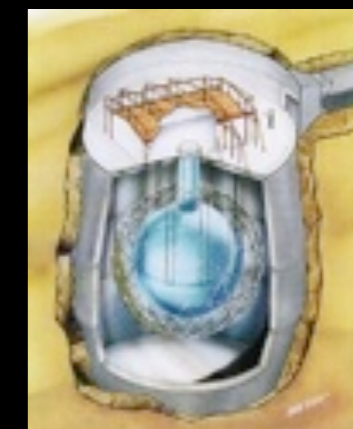


Kamland  
(3 kt)

10<sup>-43</sup>

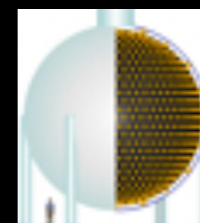
3

1



SNO (1 kt)

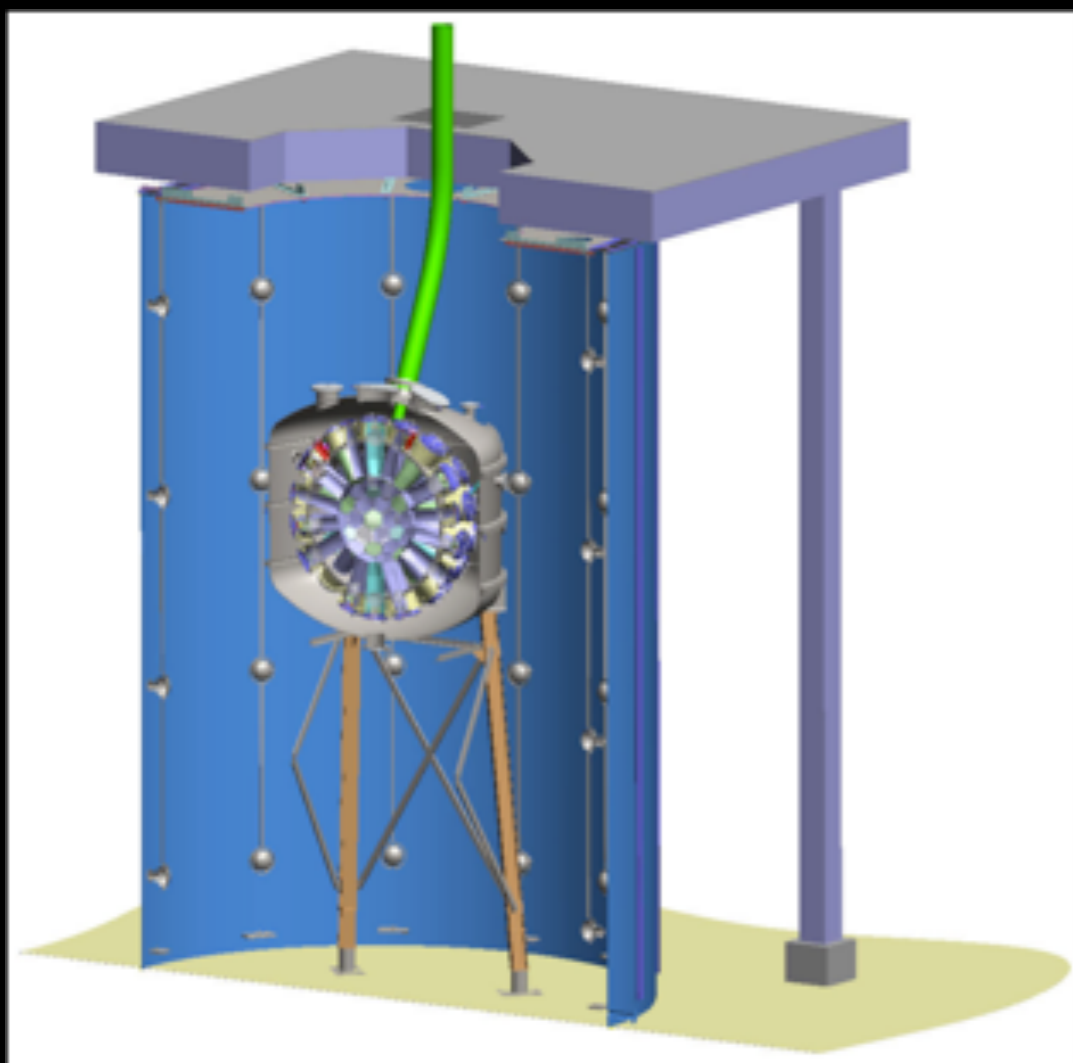
10<sup>-42</sup>



MiniBooNE  
(0.8 kt)

0.1 (0.8 kt)

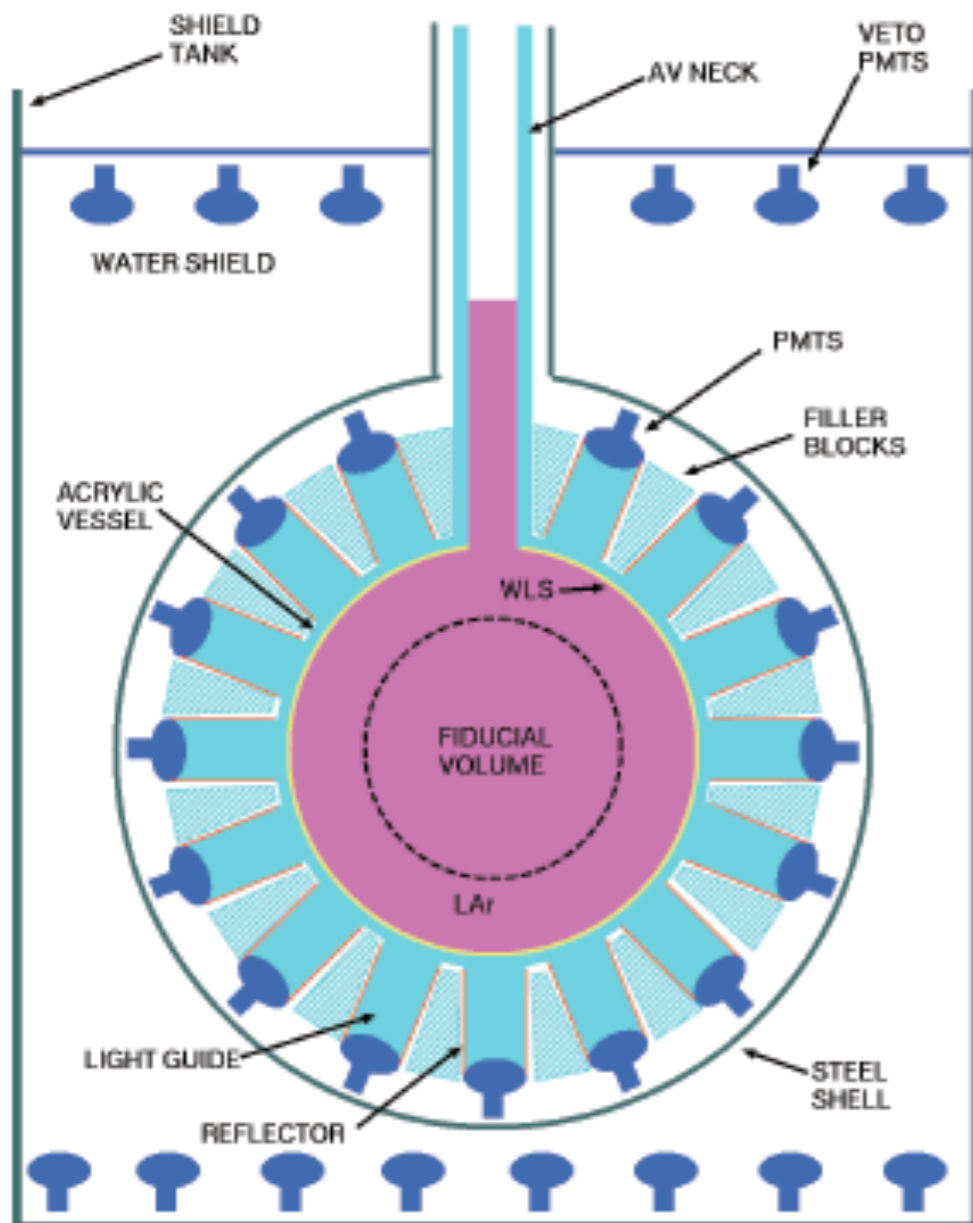
10<sup>-39</sup>



# Single Phase a la Neutrinos

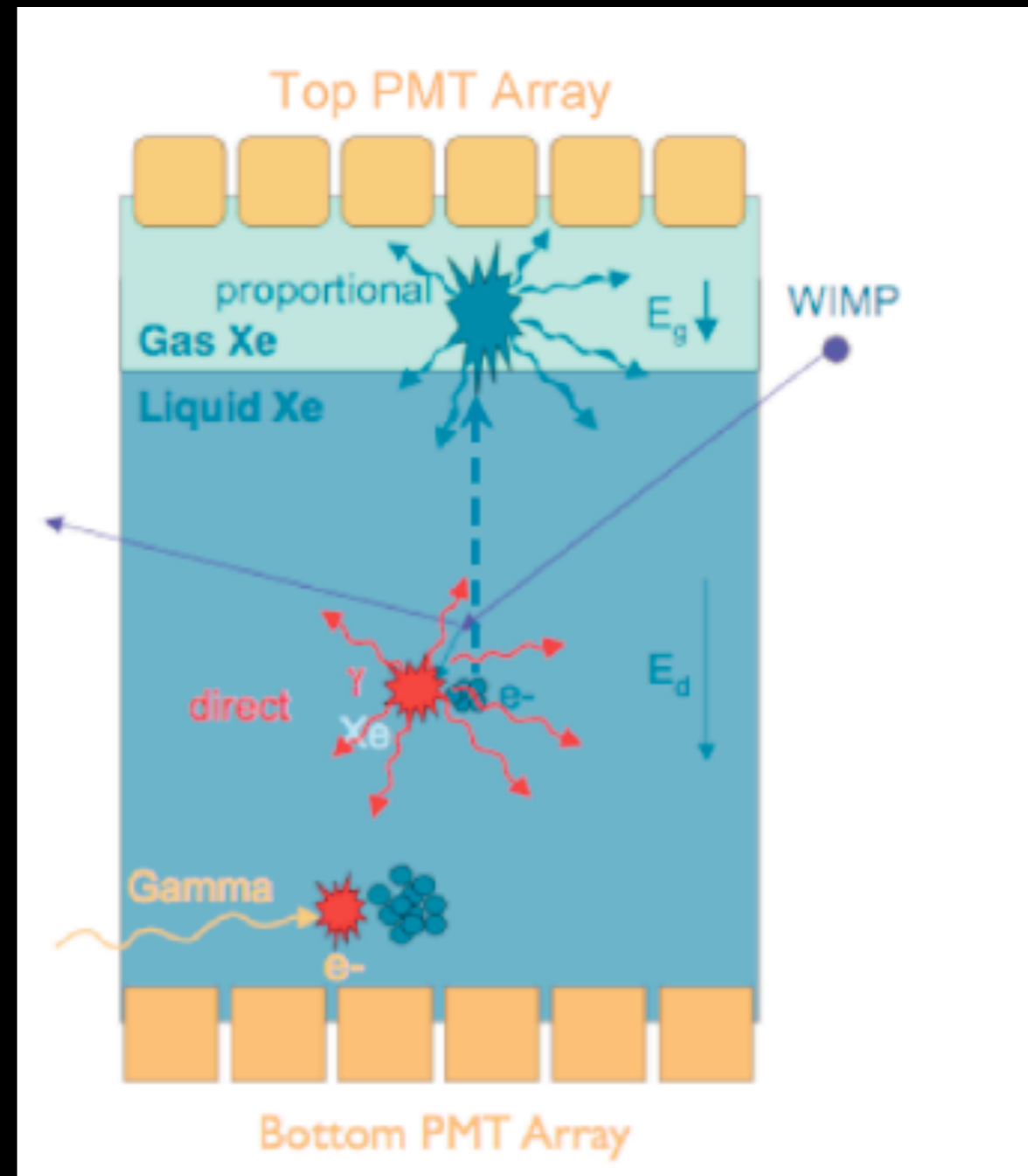
high light yield *and self-shielding* of target

background discrimination from prompt scintillation timing...



D. N. McKinsey and J. M. Doyle, J. Low Temp. Phys. 118, 153 (2000).

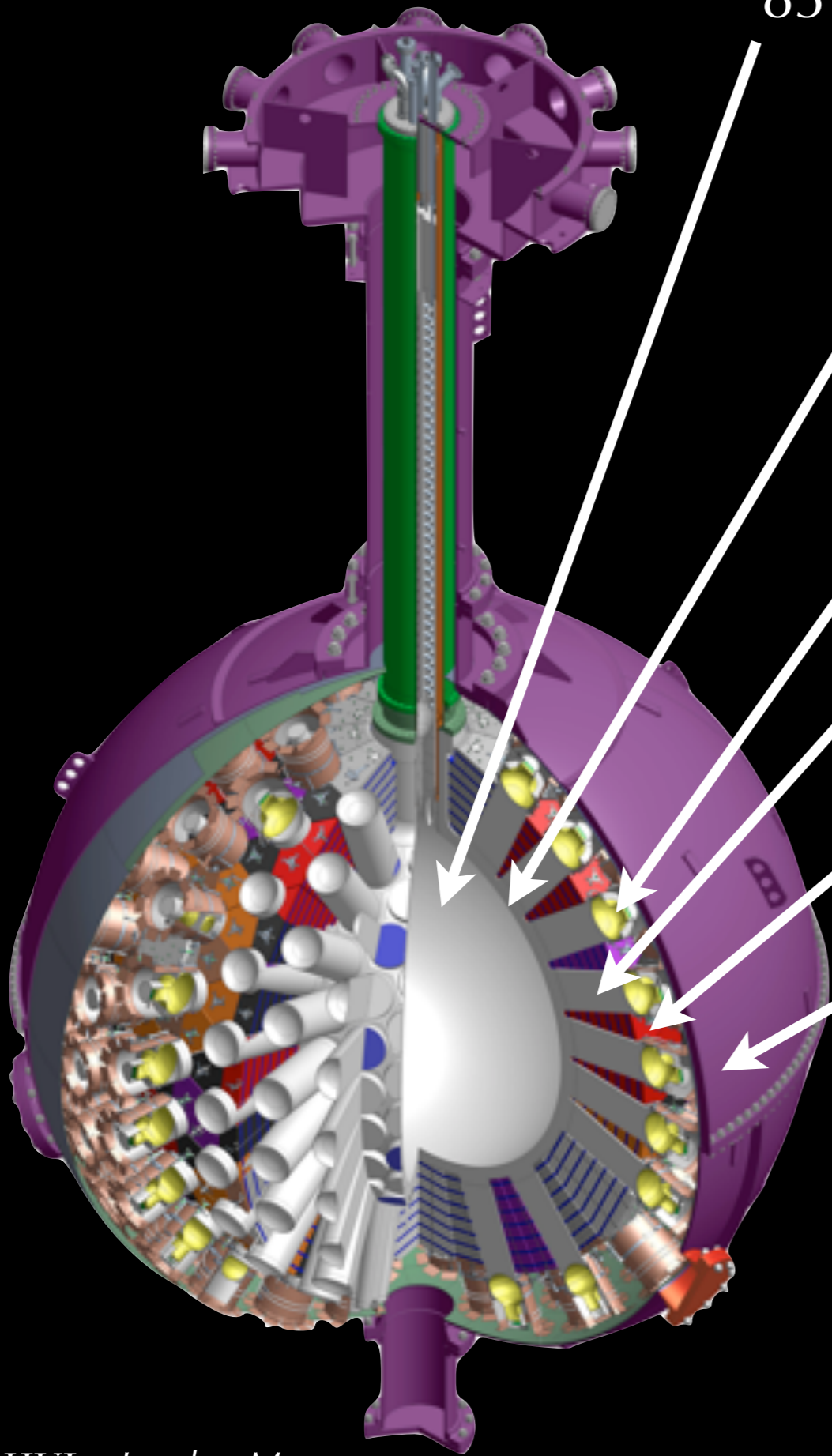
no electric fields = straightforward scalability to kT  
 1) no pile-up from ms-scale electron drift in E  
 2) no recombination in E (high photons/keVee)  
 but no charge background discrimination either!



**cf. Two Phase Detector: *and* charge**  
 (proportional scintillation)



# DEAP-3600 Detector



85 cm radius acrylic sphere contains 3600 kg of liquid argon (LAr)

1  $\mu\text{m}$  of TPB coats inside surface of sphere, wavelength shift to 420 nm

viewed by 255 8" Hamamatsu R5912 HQE PMTs (32% QE, 75% coverage)

50 cm of acrylic light guide between LAr and PMTs to mitigate PMT neutrons

PTFE filler blocks between light guides to moderate neutrons

Outer steel shell prevents LAr / water mixing (safety)

Inside 8.5 m diameter water tank, with PMTs for muon veto, and to moderate cavern neutrons and gammas

6200' underground in SNOLAB Cube Hall

# Background Strategy

## Ar-39 beta decay:

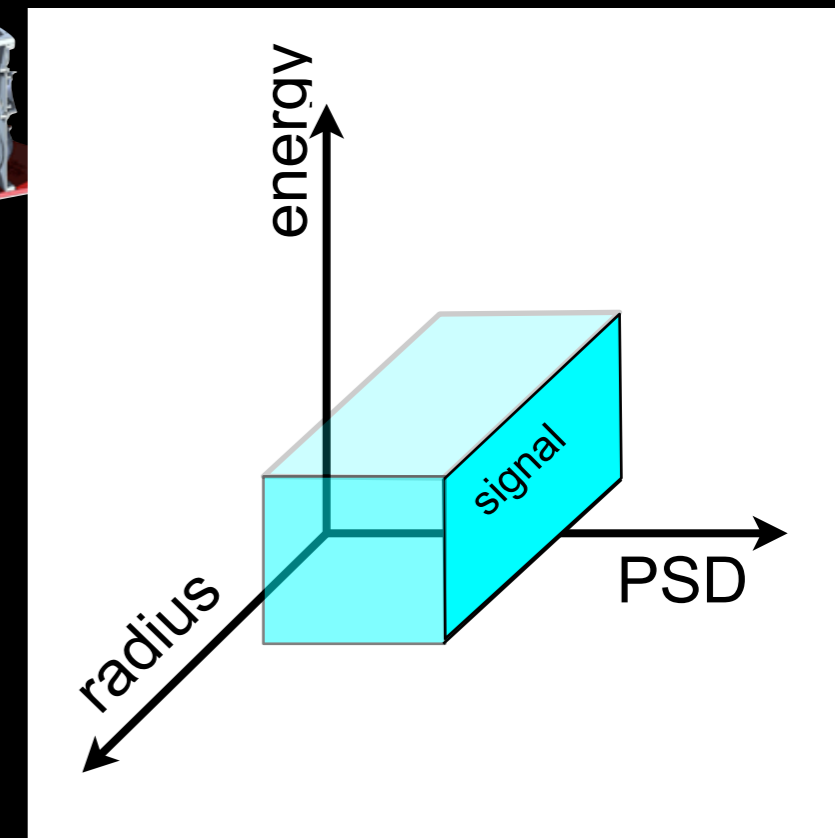
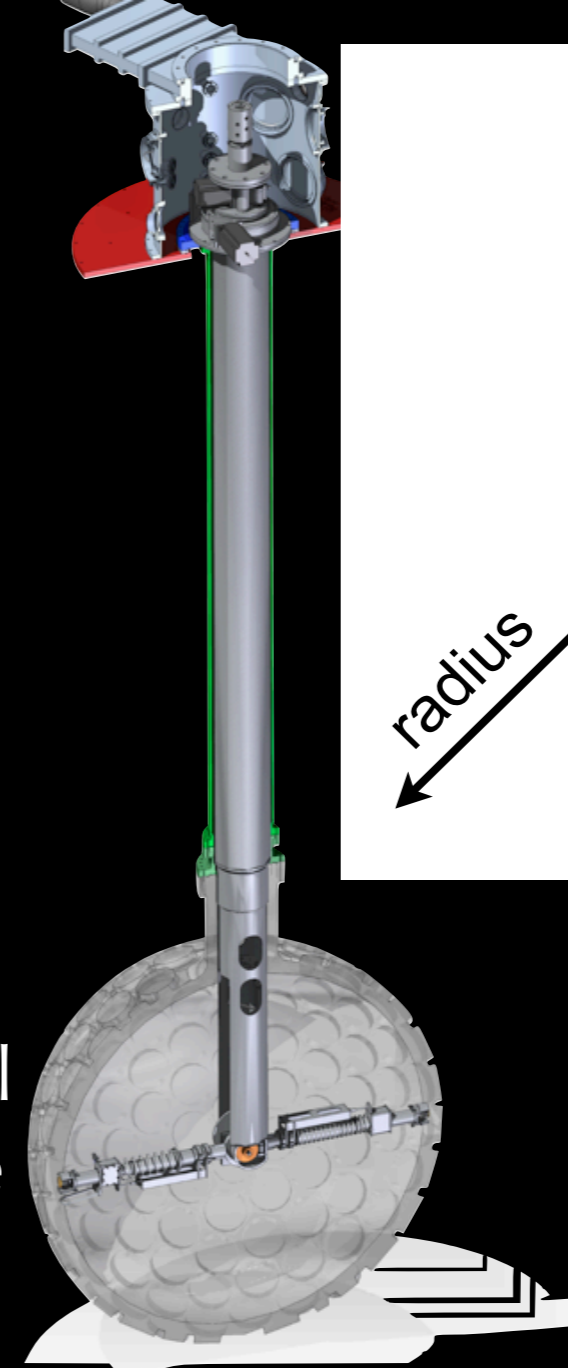
- Ar-39 decay rate  $\sim 1$  Bq/kg,  $Q=550$  keV. Dominates data rate.
- mitigated with pulse shape discrimination
- threshold for PSD determines energy threshold for dark matter search

## Alphas and Radon Progeny:

- stringent radiopurity control, ex-situ assays
- resurfacing of vessel before TPB + argon fill
- fiducialization, determines fiducial volume for dark matter search

## Neutrons and Gammas:

- passive moderation
- cross-check with active tagging: measure neutron inelastic scattering gammas
- stringent radiopurity control for (alpha,n)



Background (in Fid Vol)	DEAP-3600 Goal
Radon in Ar	$< 1.4$ nBq/kg
Surface $\alpha$ 's	$< 100$ $\mu$ Bq/m <sup>2</sup>
Neutrons (all sources)	$< 2$ pBq/kg
Ar-39	$< 2$ pBq/kg
Total (3 tonne-yr)	<b><math>&lt; 0.6</math> events</b>





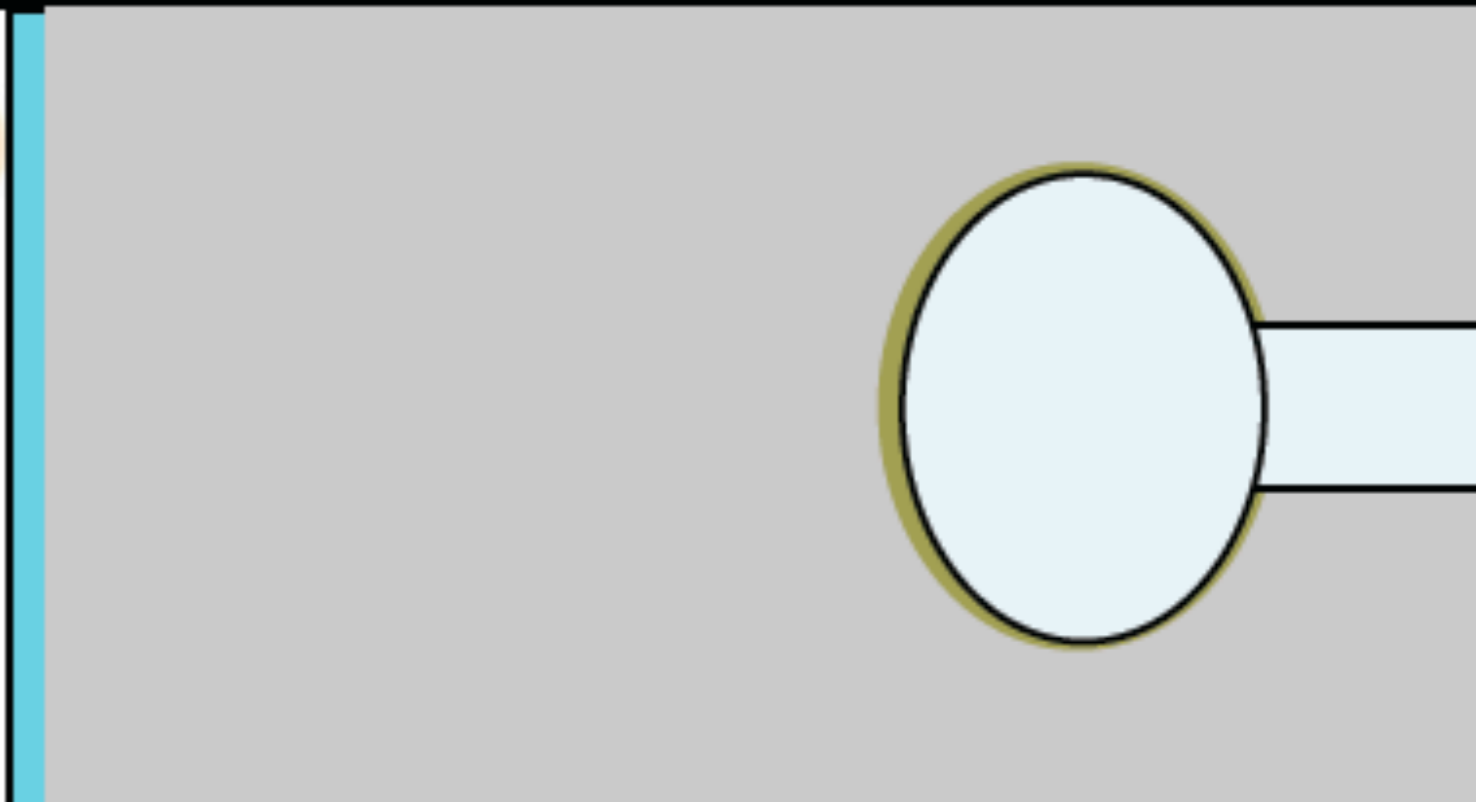
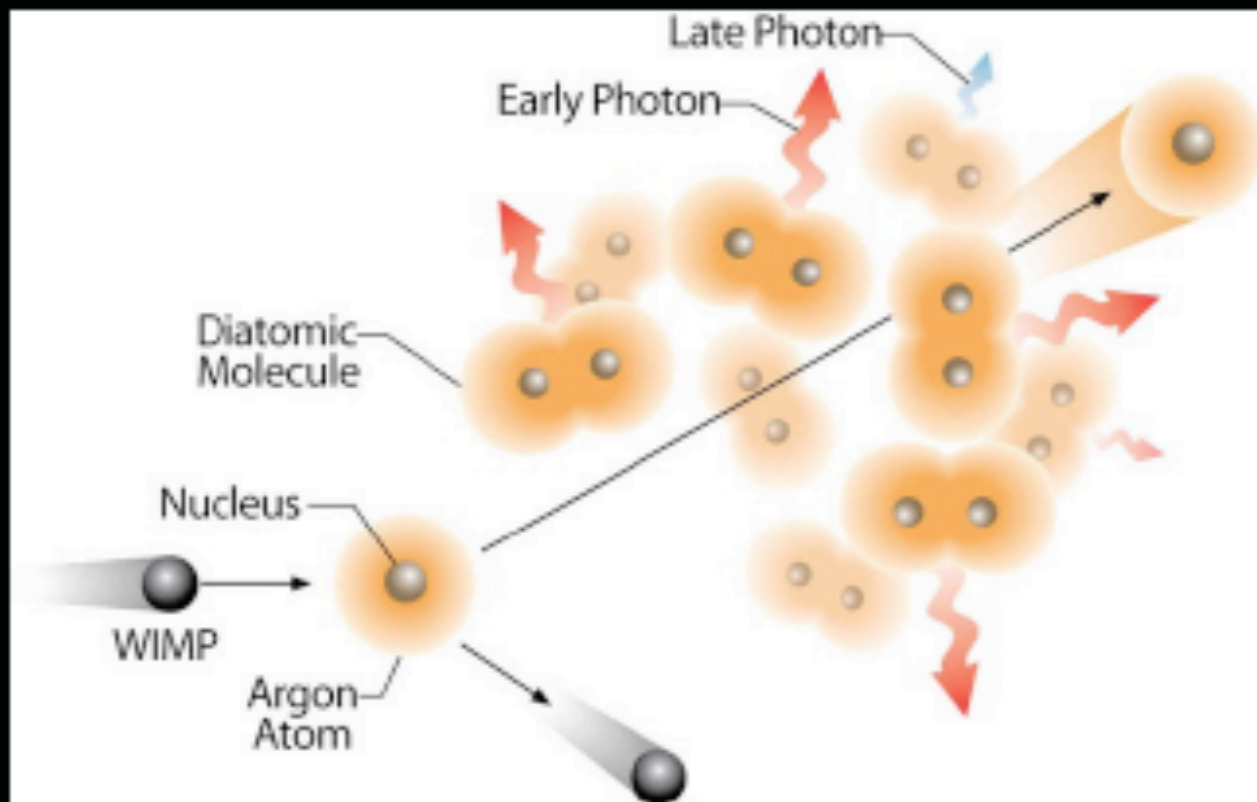
# Photon's-Eye View

EUV Light

WLS

Visible Light

Charge



Argon/Neon

TPB Acrylic/Ar/Ne

PMT

Liquid Argon dark matter target (cold! 87 K)  
LAr scintillates at 128 nm

wavelength shift  
(TPB) to  $>400$  nm

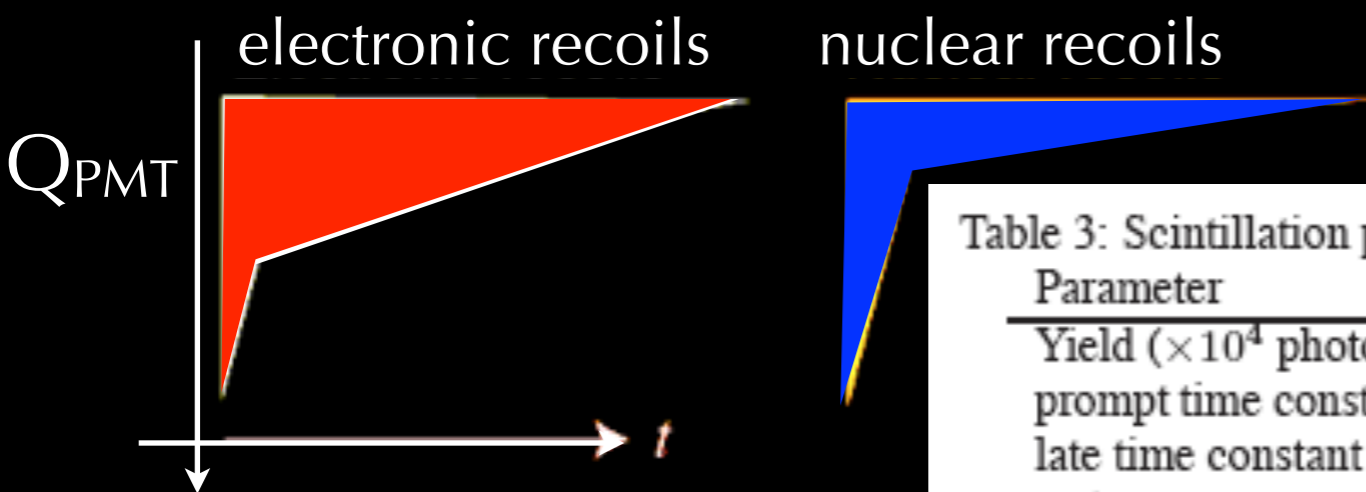
read out with PMTs,  
digitize at 250 MHz,  
maximize PE/keVee  
with  $4\pi$  coverage

*Fiducialization relies on position reconstruction from detected  $N_{PE}(t)$  in each PMT.*



# Mitigating Electron Backgrounds

LAr scintillates  $\sim 40$  photons/keV with prompt (6 ns) and slow (250x slower) components



Lippincott et al., *Phys.Rev.C* 78: 035801 (2008)

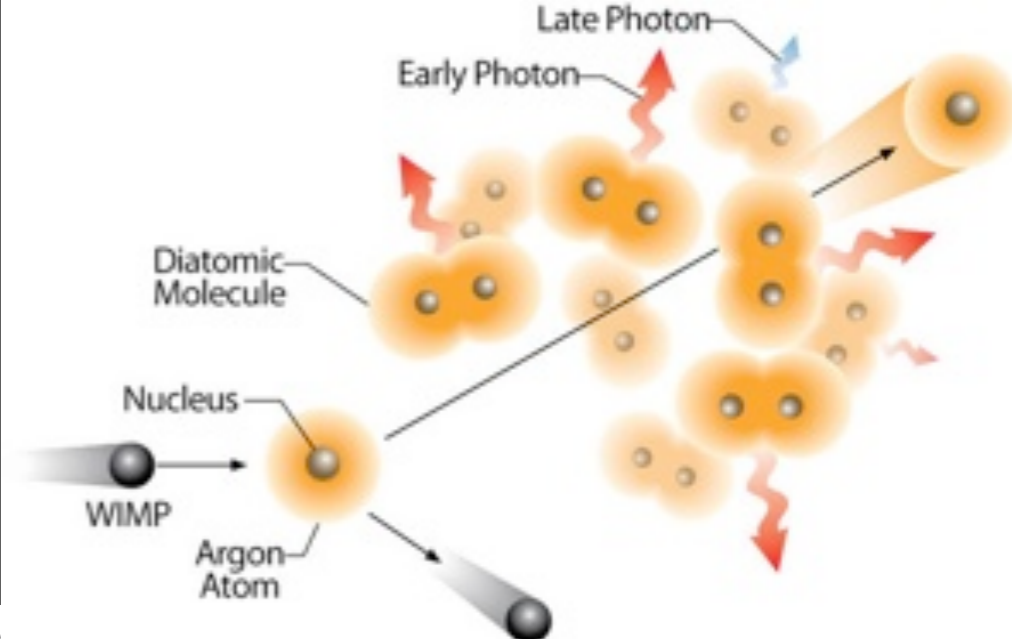
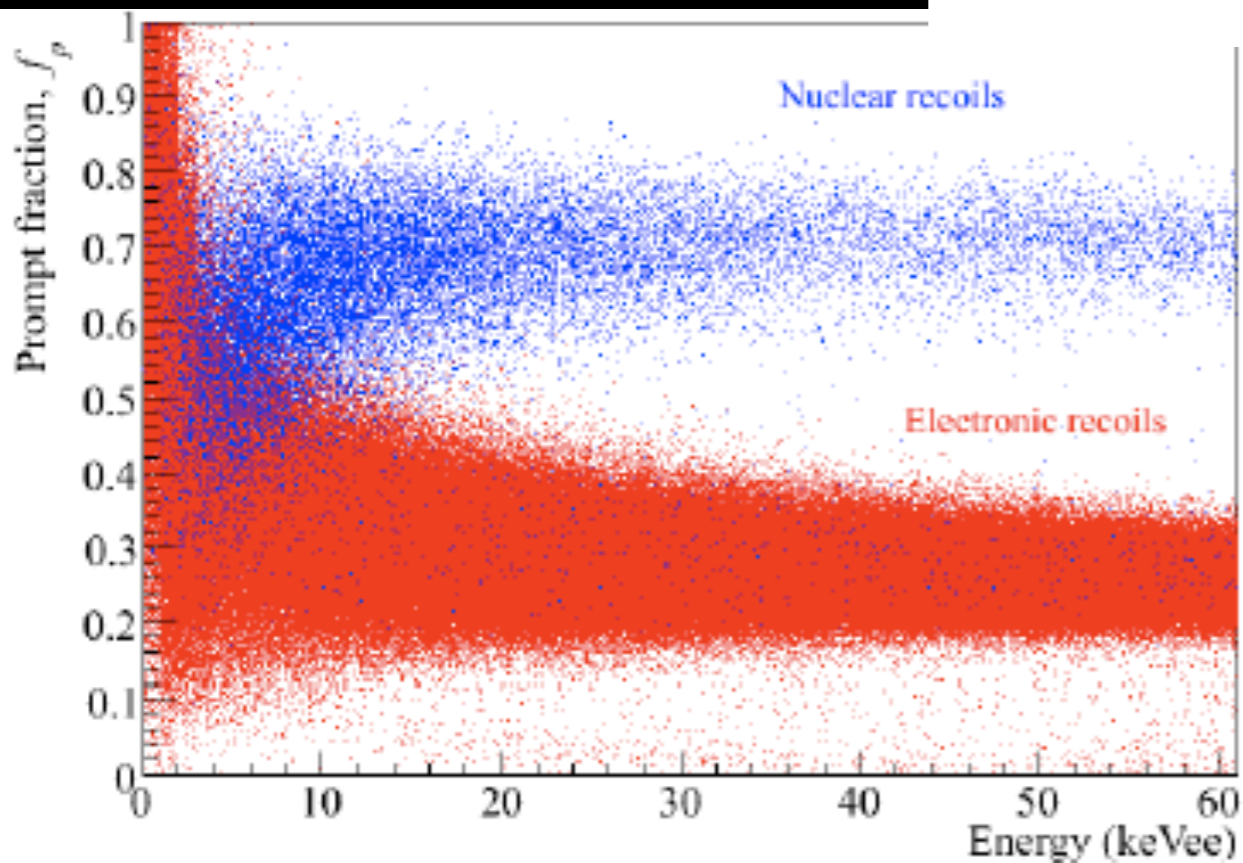


Table 3: Scintillation parameters for liquid neon, argon, and xenon.

Parameter	Ne	Ar	Xe
Yield ( $\times 10^4$ photons/MeV)	1.5	4.0	4.2
prompt time constant $\tau_1$ (ns)	2.2	6	2.2
late time constant $\tau_3$	$15 \mu\text{s}$	$1.59 \mu\text{s}$	21 ns
$I_1/I_3$ for electrons	0.12	0.3	0.3
$I_1/I_3$ for nuclear recoils	0.56	3	1.6
$\lambda(\text{peak})$ (nm)	77	128	174
Rayleigh scattering length (cm)	60	90	30



identify, reject electronic backgrounds via pulse shape vs. time difference

McKinsey & Coakley, *Astropart. Phys.* 22, 355 (2005).  
Boulay and Hime, *Astropart. Phys.* 25, 179 (2006)

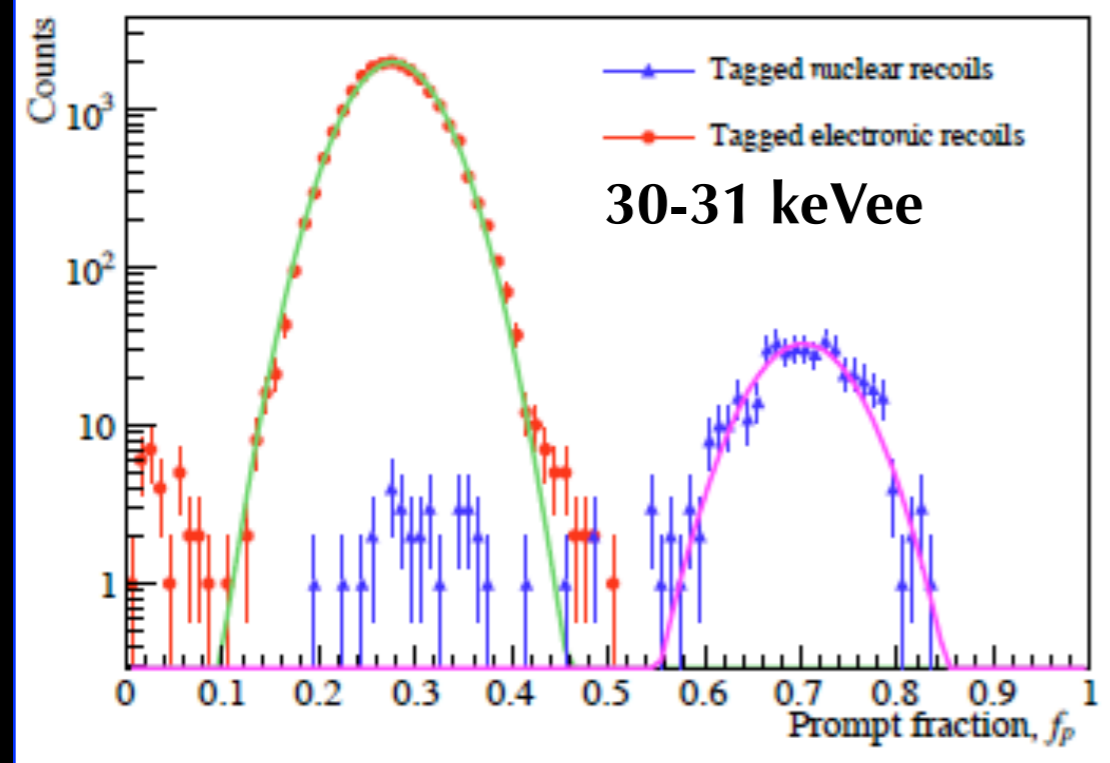
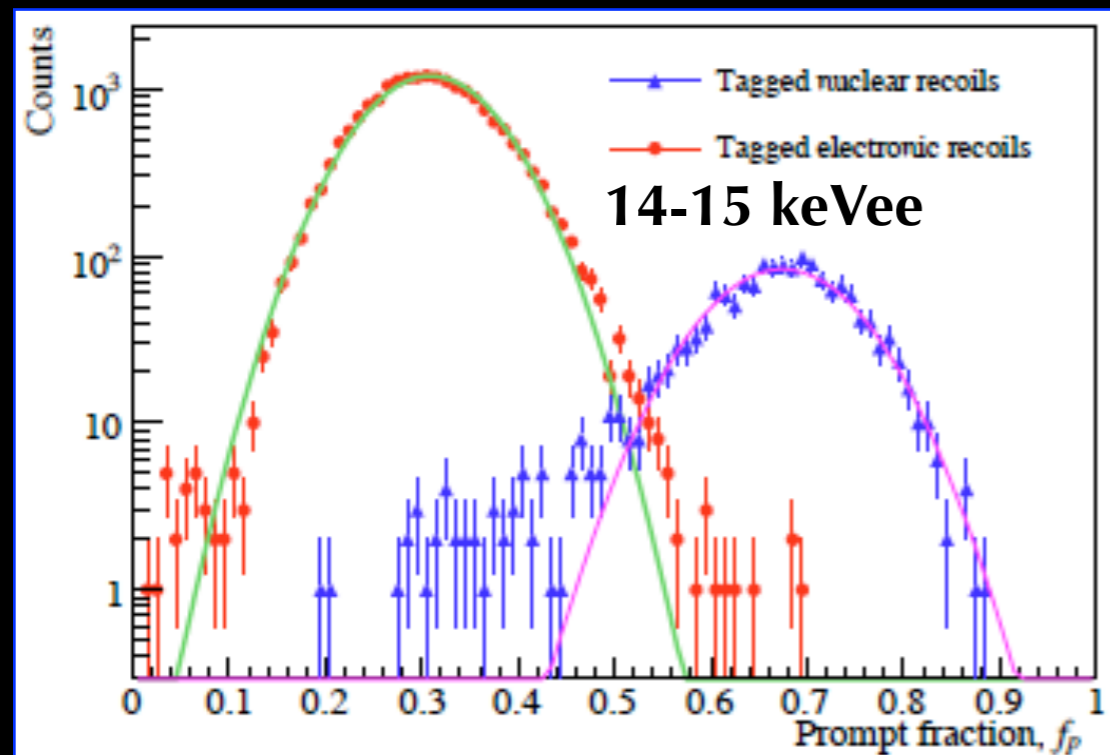
Critically important for LAr:

Ar-39 beta (1 Bq/kg)!

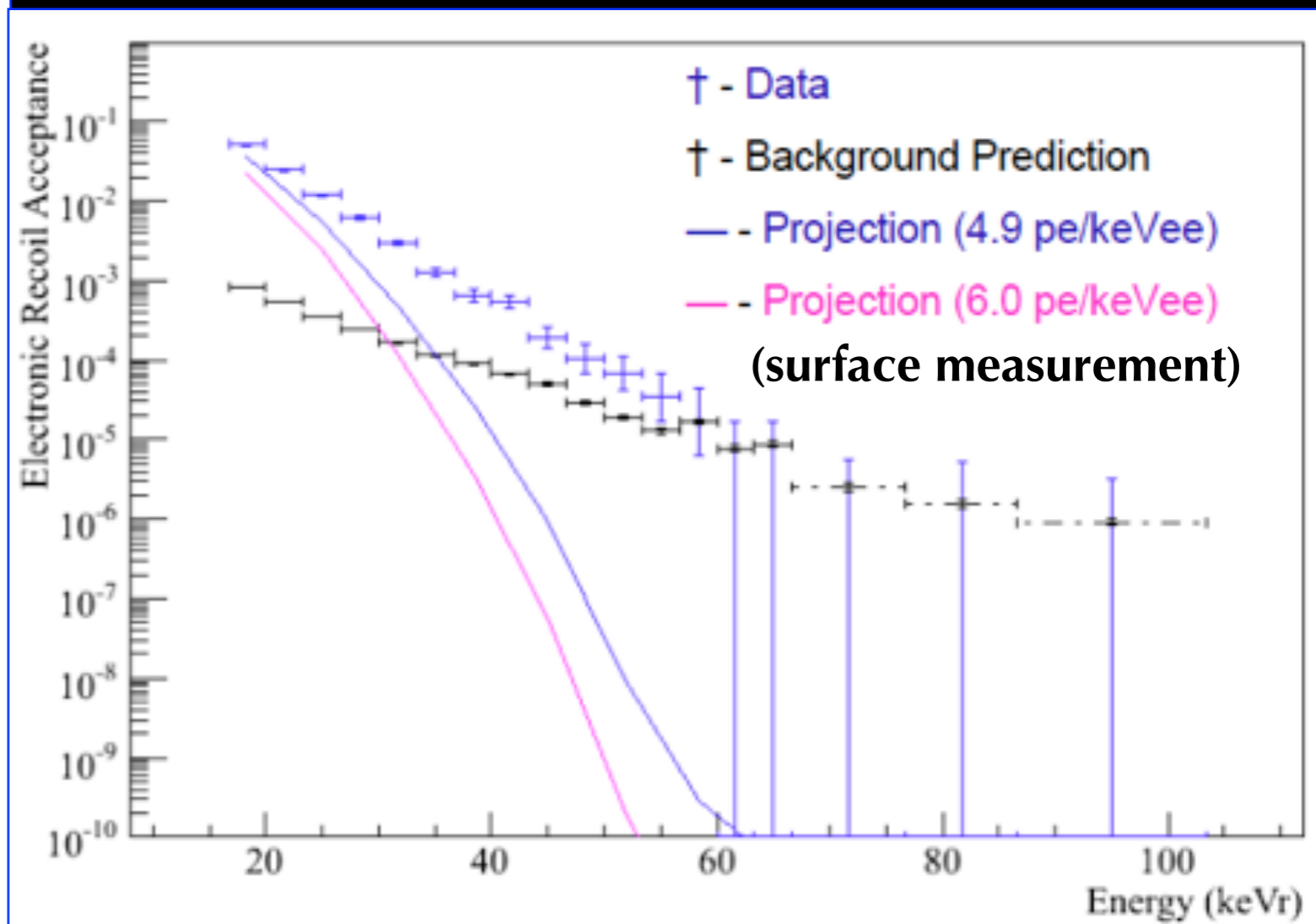


# Pulse Shape Discrimination

Fraction of prompt light discriminates between electronic and nuclear recoils.



*Leakage depends strongly on light yield!*



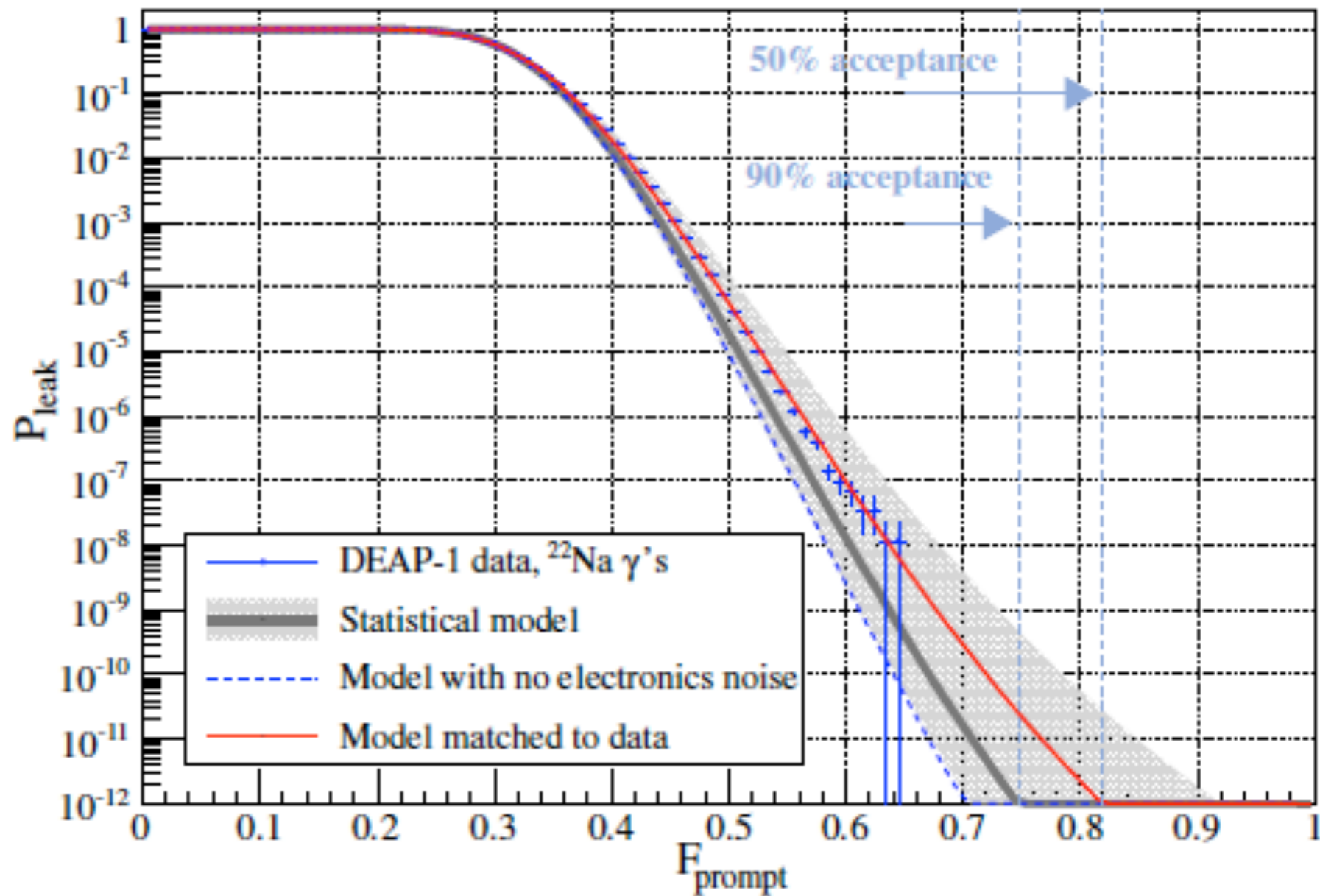
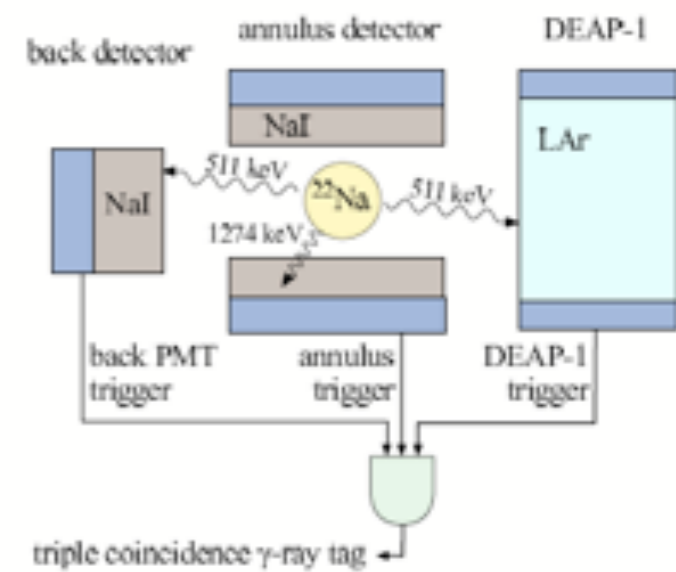
Single-phase LAr detectors possible because of rejection power from timing alone: in principle, potential for kT scale detectors. *Boulay and Hime, Astropart. Phys. 25, 179 (2006)*

# High Statistics PSD

high intensity tagged Na-22 gamma source with DEAP1  
prototype detector, underground at SNOLAB  
integrated 1.2E8 tagged gammas  
detector light yield:  
 $2.8 \pm 0.1$  PE/keVee

PSD leakage:  
<  $2.8 \times 10^{-8}$  @ 90% recoil  
acceptance,  
in 120-240 PE  
(45-88 keVee in DEAP1)  
*Amaudruz et al., arXiv:09*

simple model predicts  
 $O(10^{-10})$  leakage @ 50%  
recoil acceptance in  
120-240 PE in  
DEAP-3600 with  
8 PE/keVee light yield





# PSD with Other Statistics

Variance of PE counting significantly affects pulse shape discrimination

Developed new Bayesian PE-counting method which reduces variance substantially:

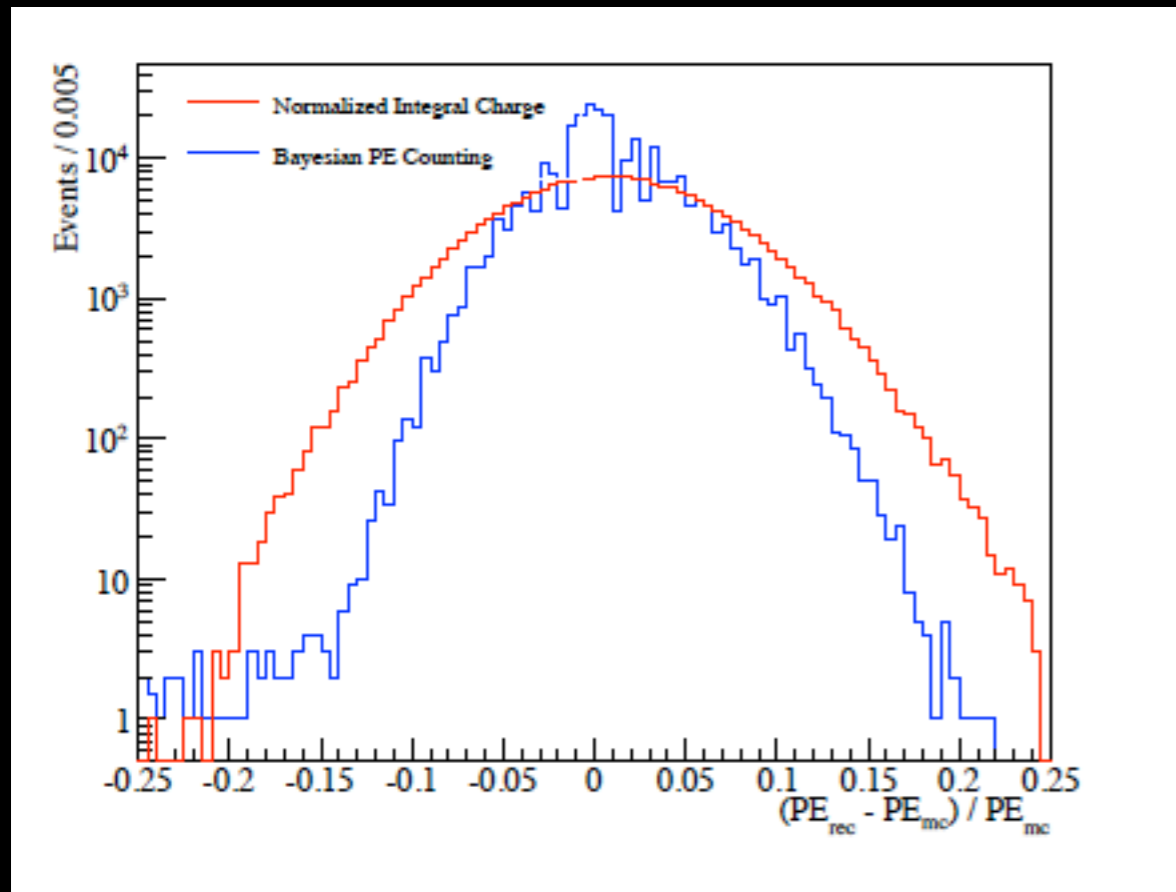


Figure 12: The ratio of estimated number of photoelectrons divided by the true number of photoelectrons for simulated  $^{39}\text{Ar}$  decays in the MiniCLEAN. Only events between 75 and 100 PE, and radius less than 295 mm are shown.

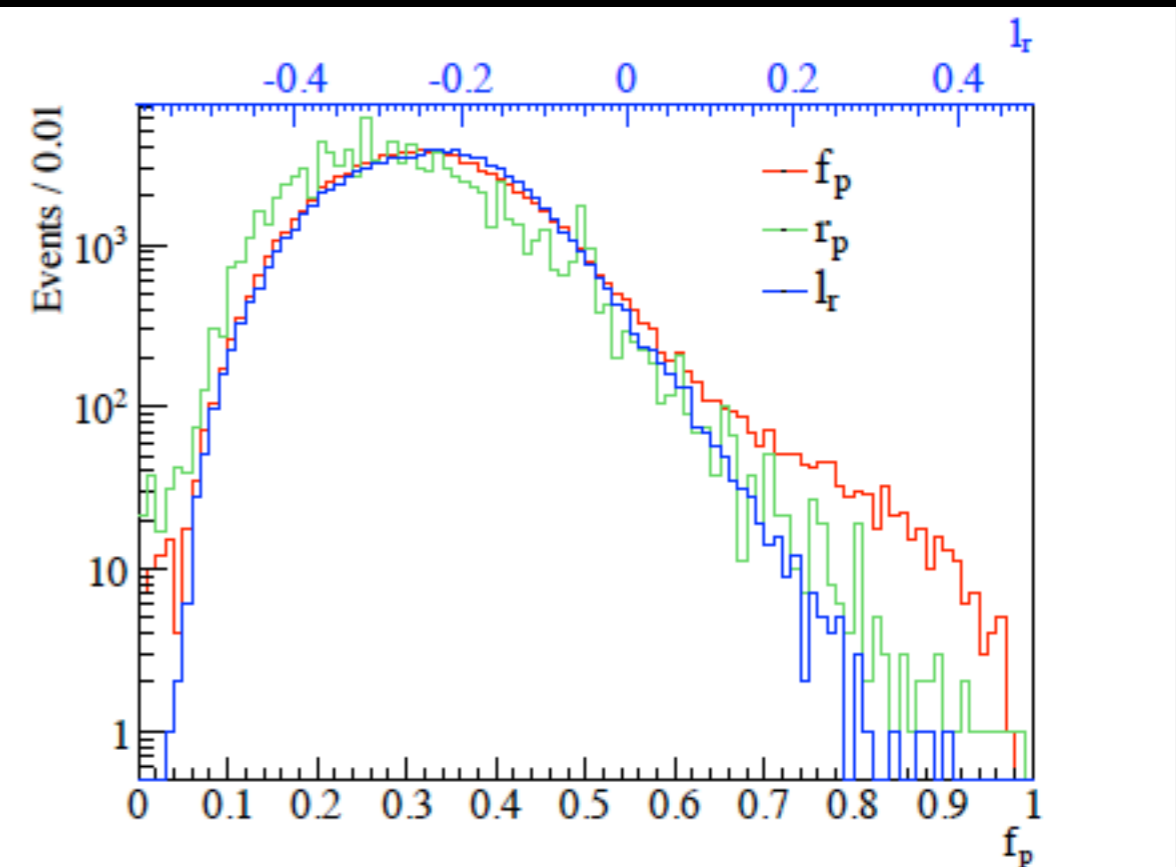


Figure 8: Distribution of  $f_p$ ,  $r_p$ , and  $l_r$  test statistics for electronic recoils for  $^{22}\text{Na}$  calibration events in DEAP-1 with 30 PE. The  $l_r$  values have been linearly transformed such that the median values for the electron and nuclear recoil distributions match those for  $f_p$ . The shift in the  $r_p$  peak relative to  $f_p$  is due to the discrete nature of the test statistic.

PSD with:

- (a) fraction of prompt Q,
- (b) fraction of prompt PE
- (c) Lrecoil: likelihood ratio test of e- vs. nuclear recoil hypothesis

7x reduction in e- tail above  $f_p = 0.7$

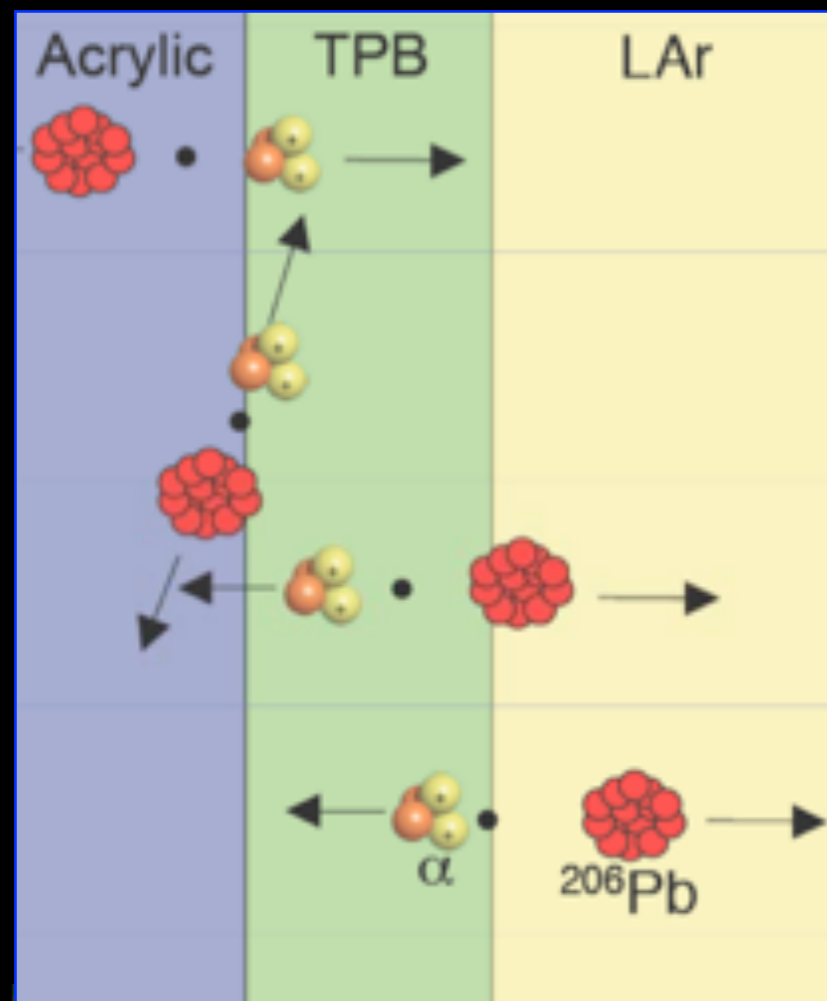


# Mitigating Alpha Backgrounds

Dangerous Radon (Rn) backgrounds from decay of Rn progeny on surfaces between Acrylic Vessel (AV) and TPB, and in TPB.

Dominant source of Rn from plate-out on AV during manufacture and construction.

So, sand off 1 mm of acrylic from inside of AV before TPB deposition, x25 reduction.



*With a robot!*

Resurfacer testing: (movie) complete this week, start resurfacing actual AV next week!



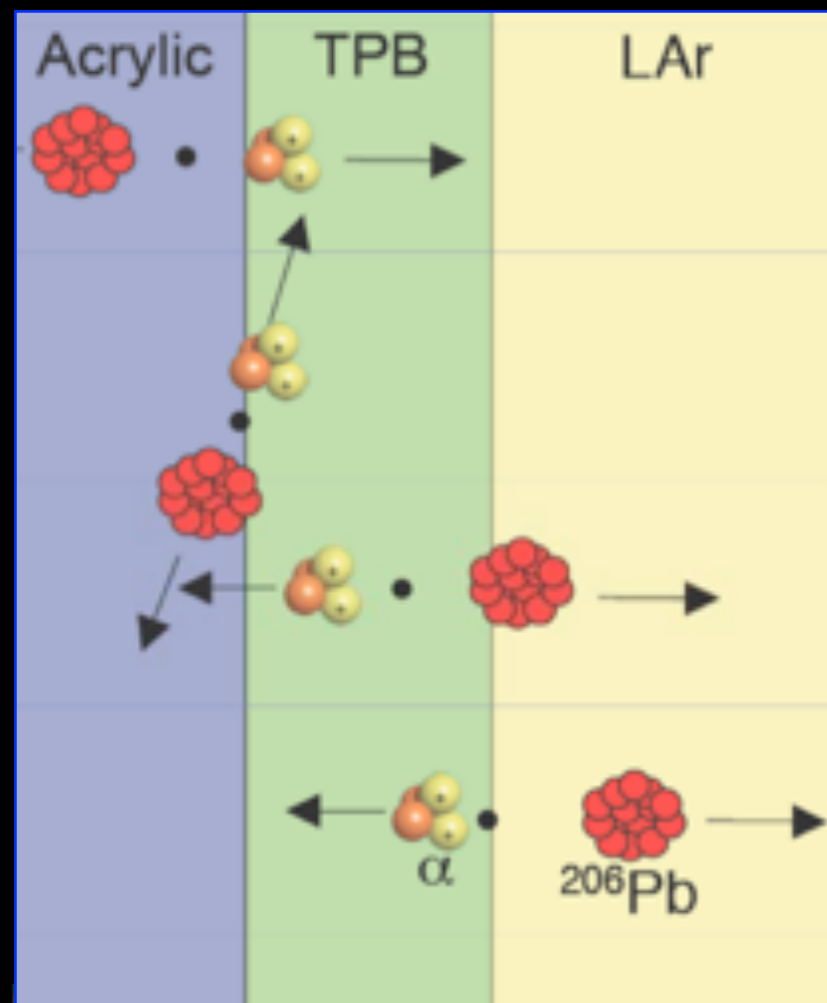


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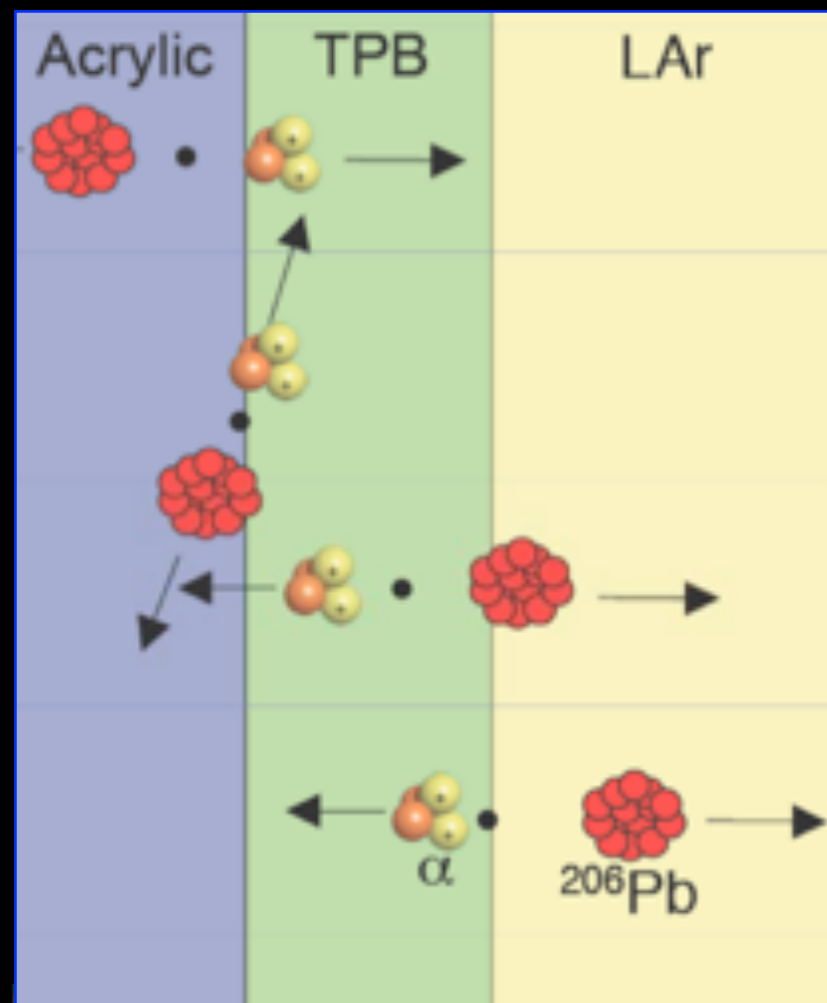


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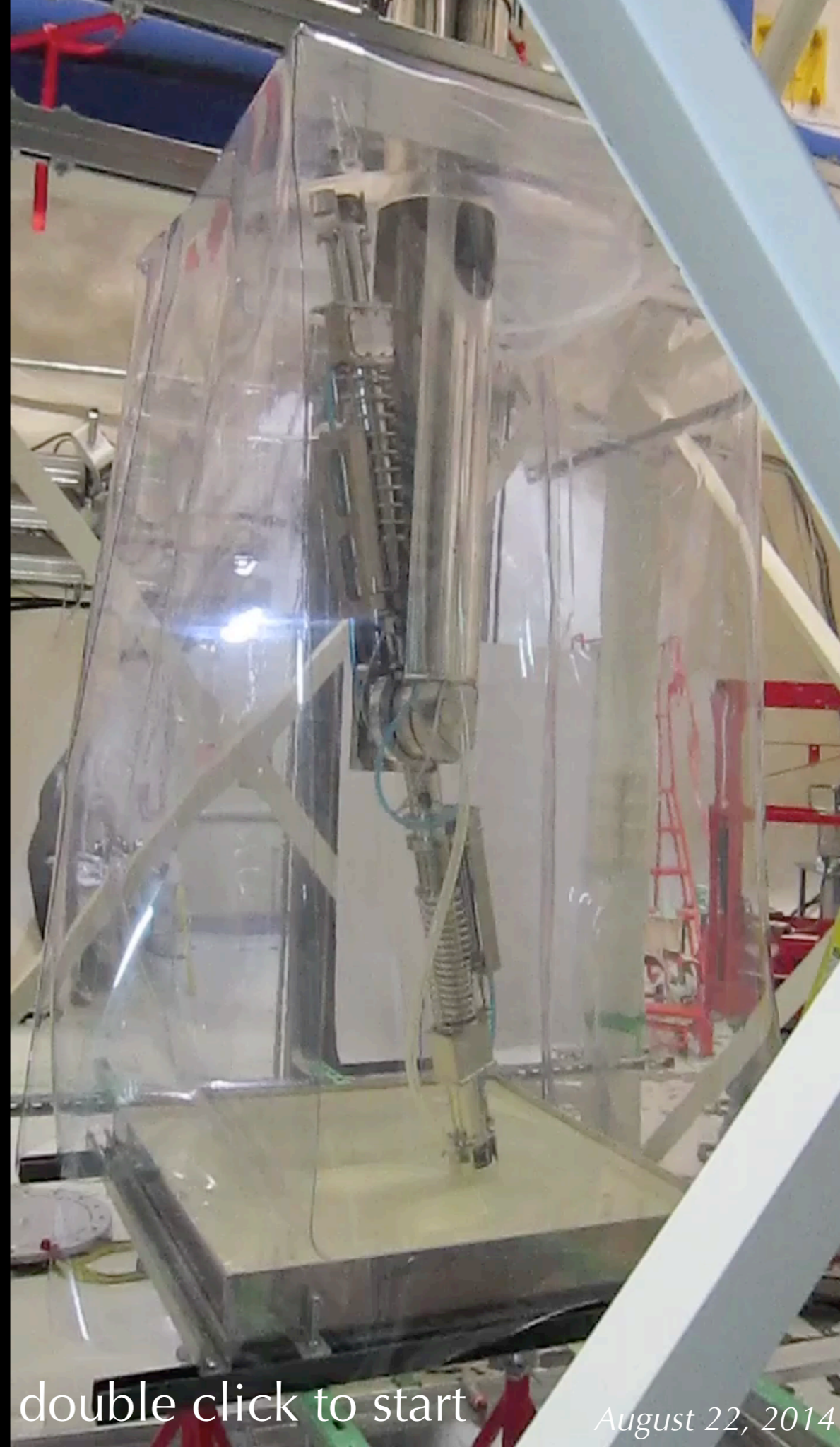
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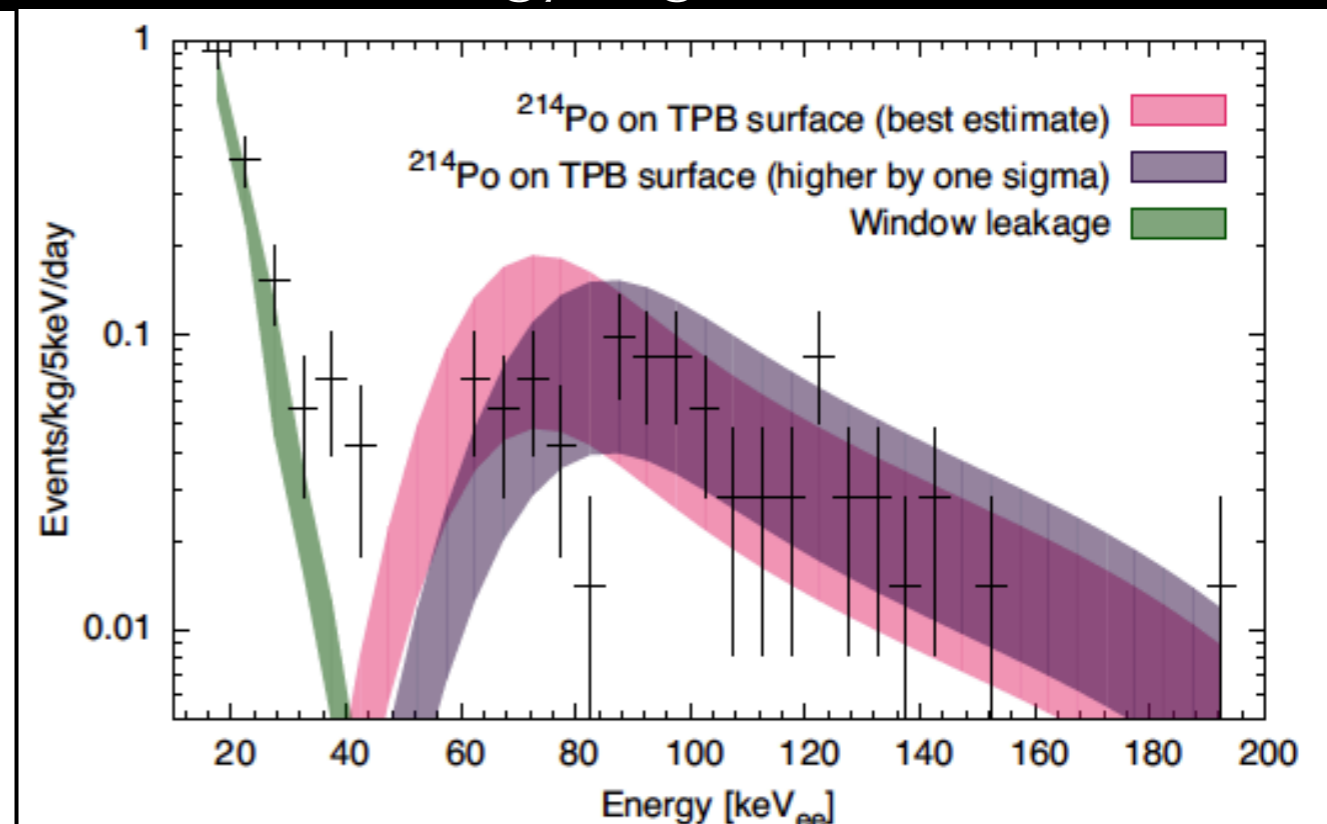
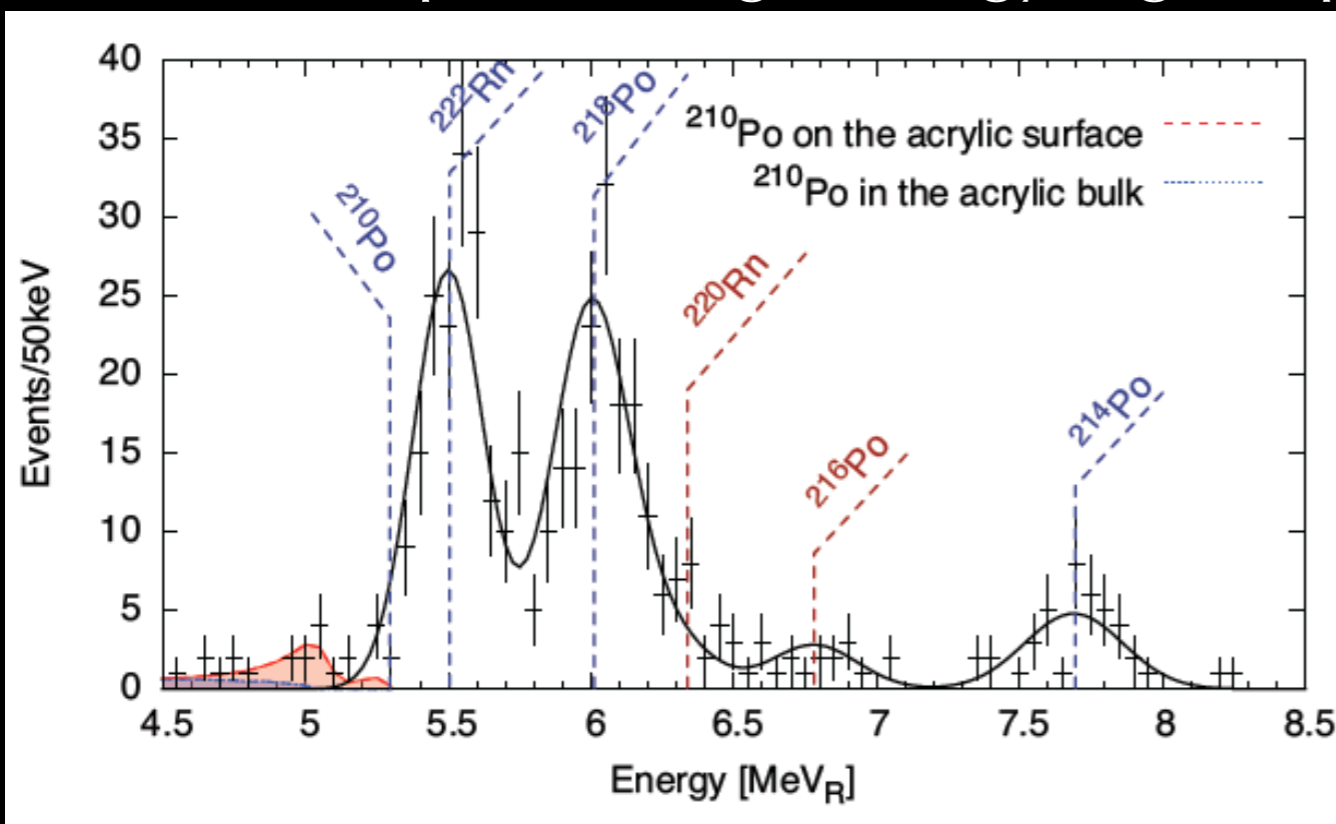
double click to start

August 22, 2014



Main tools for controlling alpha backgrounds remaining after resurfacing are:

1. position reconstruction: fiducial volume cut defined on reconstructed radius
    - leakage determined by position reconstruction resolution
    - cut set for  $<0.6$  events in 3 years run, results in 1 Tonne LAr fiducial mass
  2. Radiopurity control: substantial R&D in DEAP-1 prototype detector(s):
    - 16 uBq/kg background in energy region of interest well understood, described by 1) Rn-decay on surfaces, and 3) PSD leakage at fiducial edge
- measure alphas in high-energy region, predict low energy region (120-240 PE)



# Alpha Scintillation in TPB

TPB wavelength-shifts from 128 nm to visible (fluorescence)  
ex-situ test benches for spectrum, efficiency, angular dist.

*V. M. Gehman et al., NIM A 654 1 (2011) 116-121*

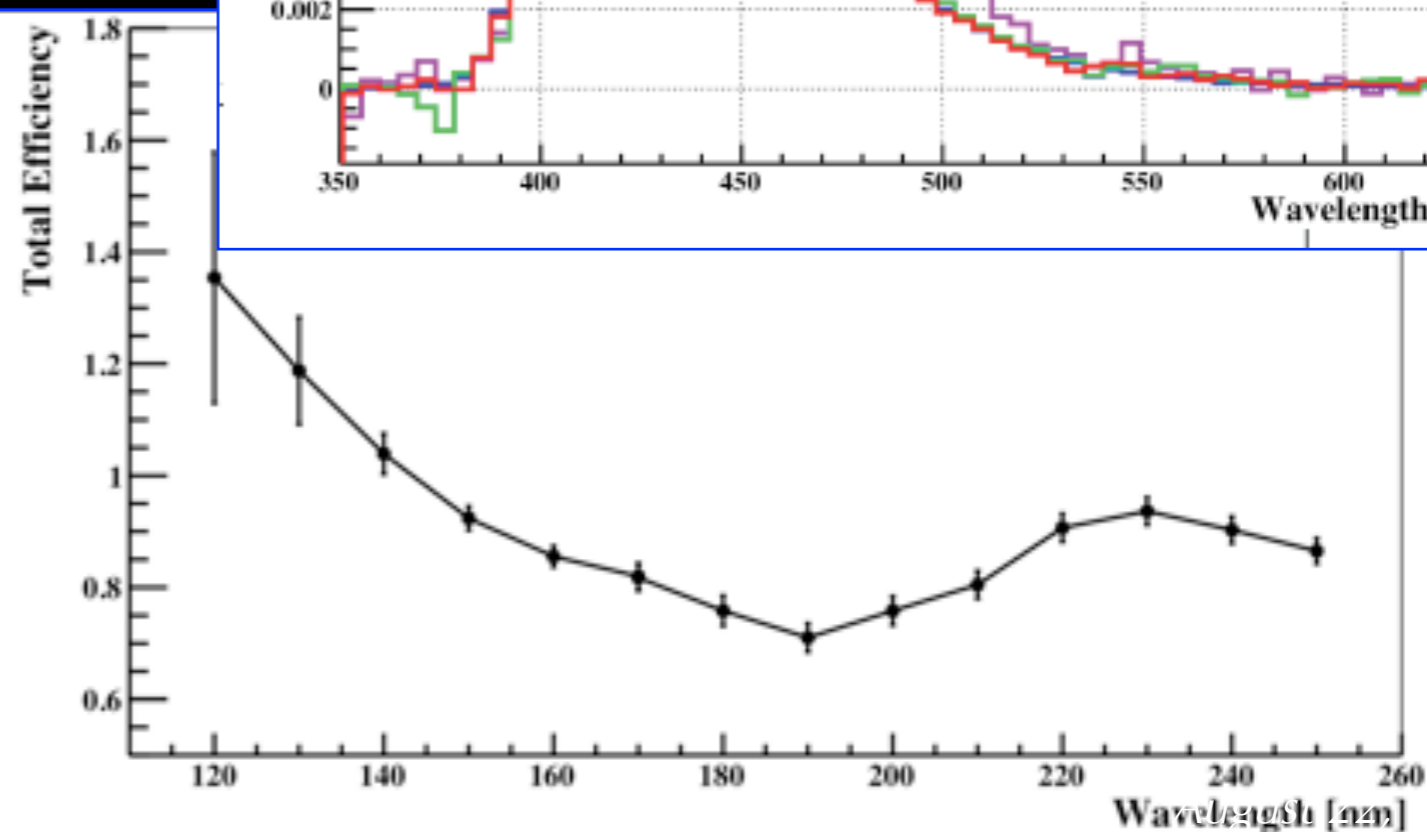
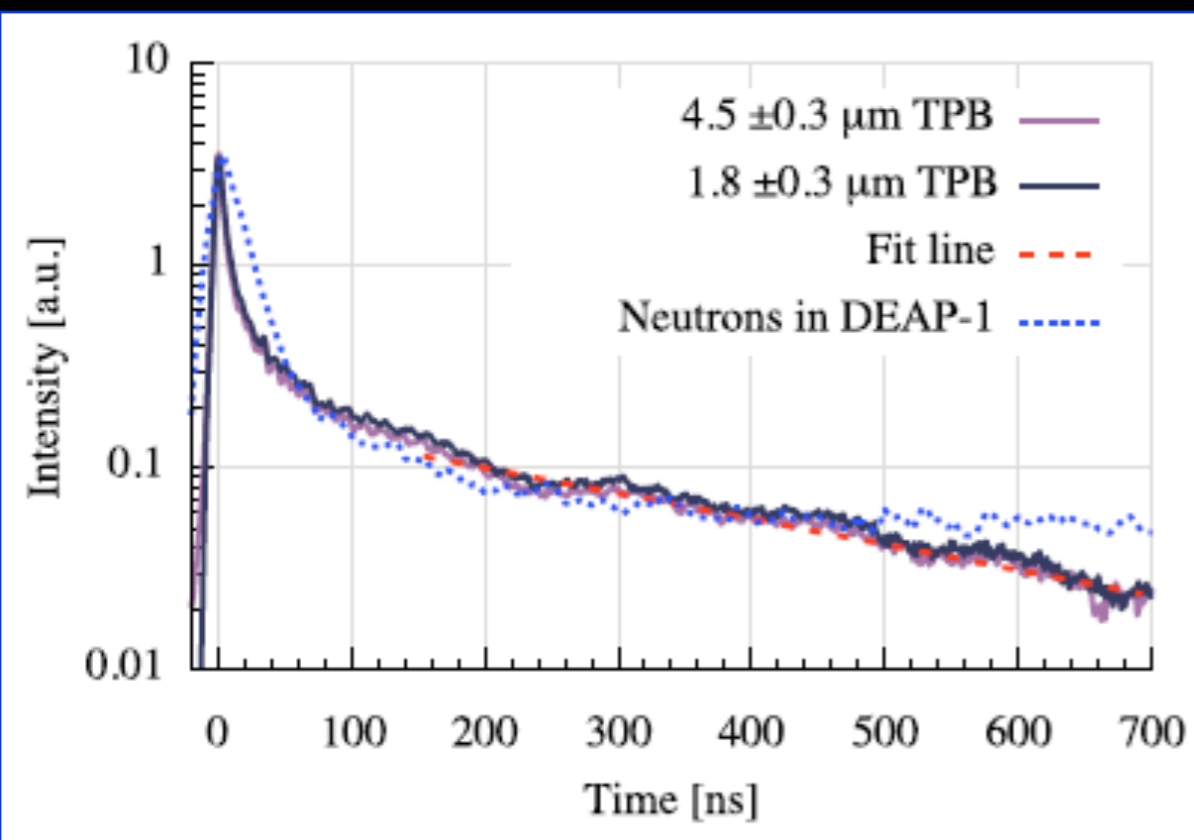
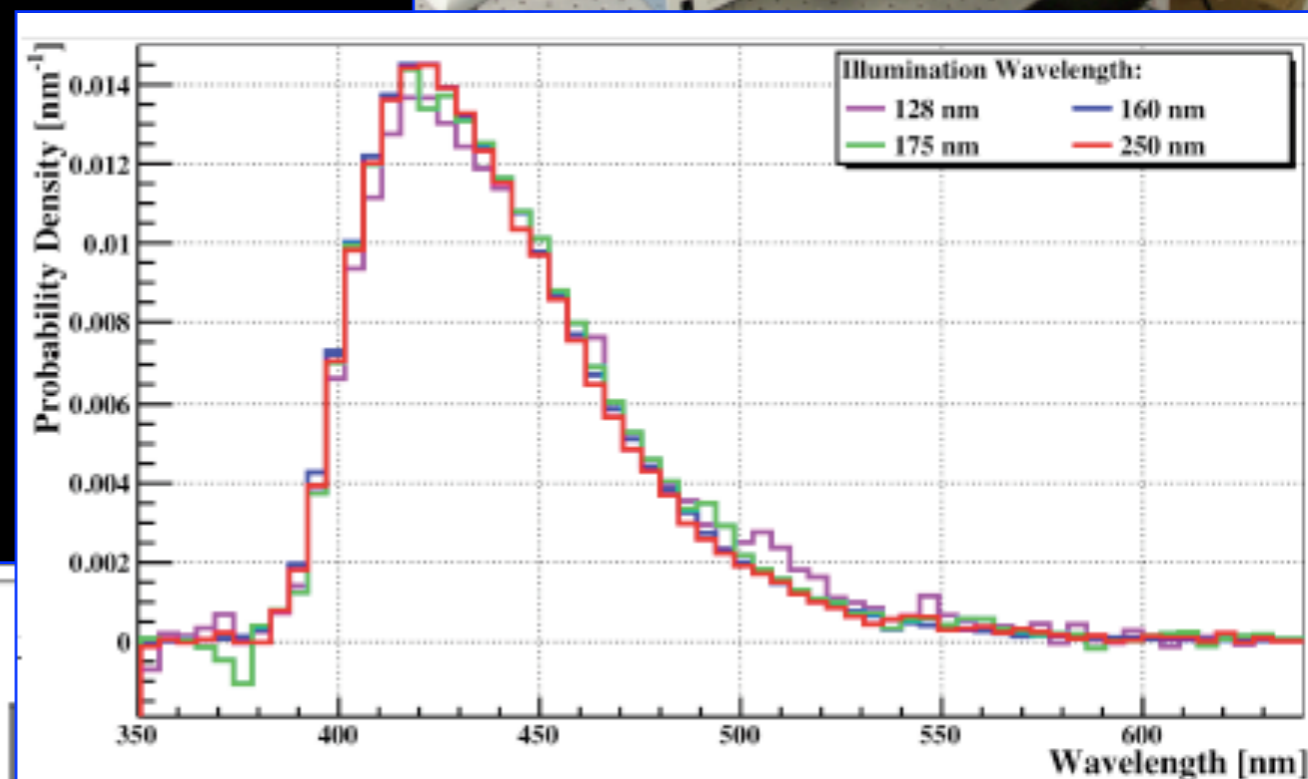
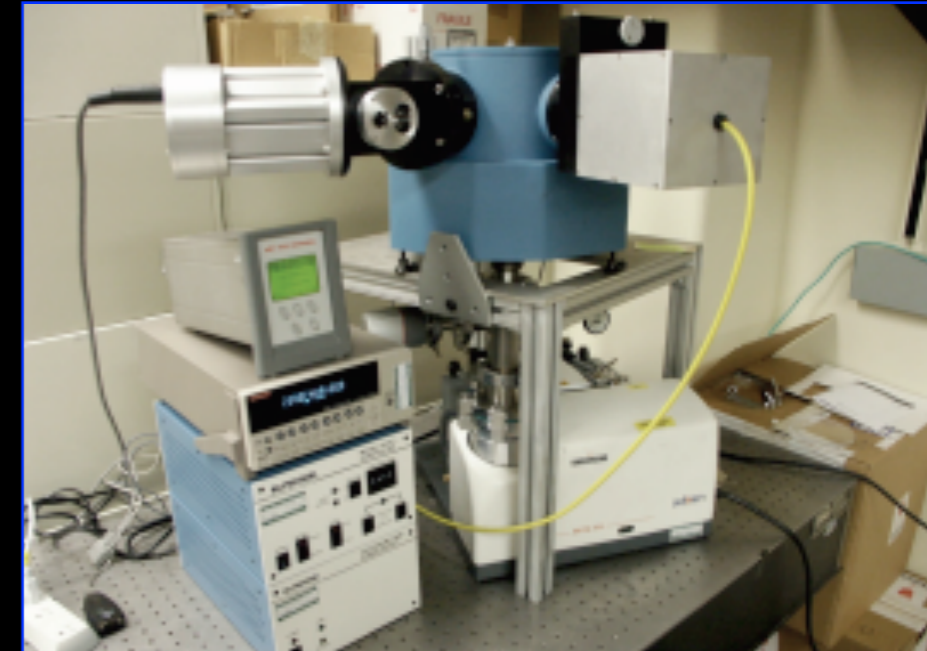
Alpha scintillation in TPB has rejection power,

ex-situ test stand finds  $11 \pm 5$  and  $275 \pm 10$  ns  
fast and slow time constants, and fast:total

intensity ratio of  $0.67 \pm 0.03$

(cf. 7 ns and 1600 ns, and 0.75)

*T. Pollmann et al., NIM A 635 1 (2011) 127-130*



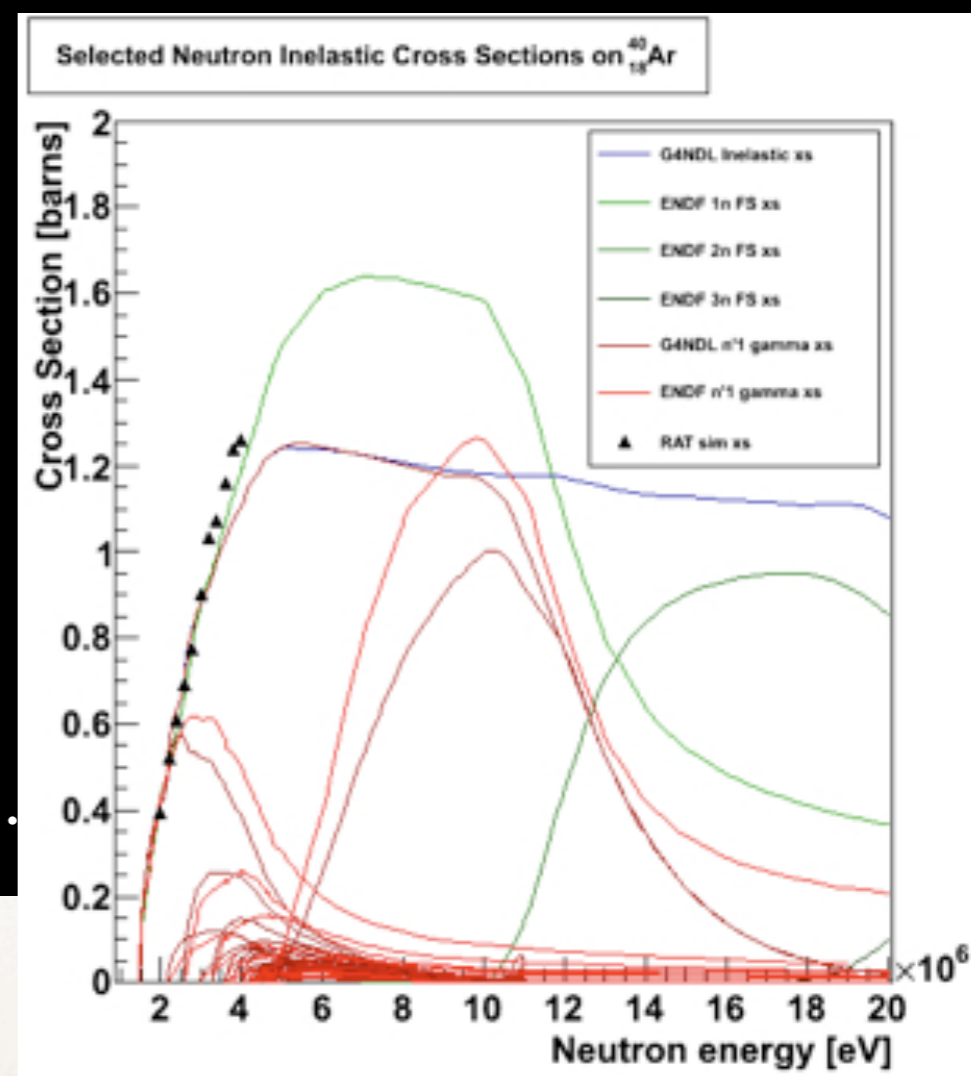


# Mitigating Neutron Backgrounds

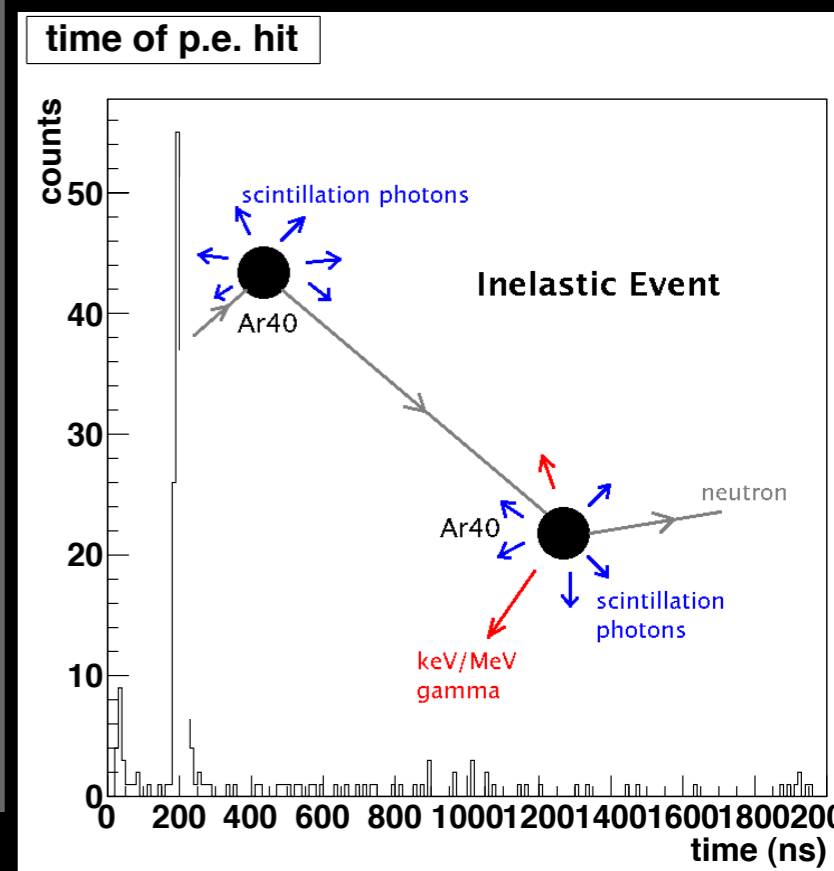
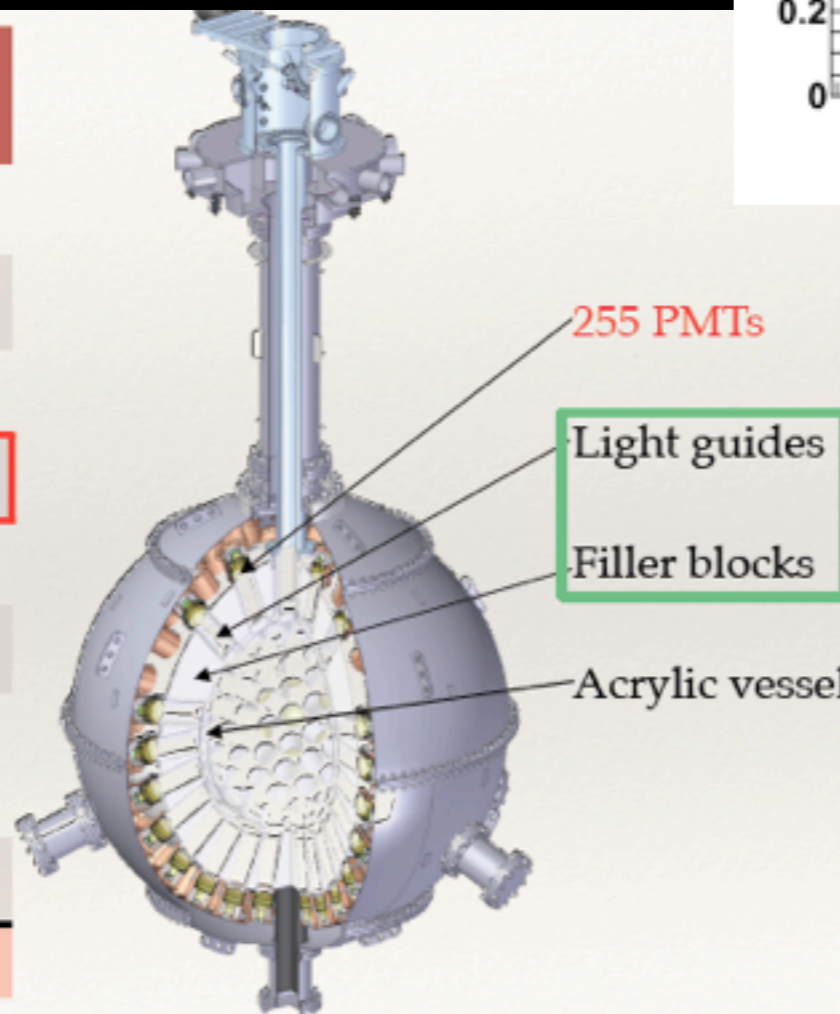
Dominant source of neutron backgrounds comes from (alpha,n) in PMT glass.

So, shield LAr target from PMTs by 50 cm of acrylic to moderate this neutron flux.

Cross-check n attenuation using external tagged AmBe source, tag internal n using inelastic gammas.



(In 3 years)	# of neutrons (produced)	Events in ROI
Acrylic vessel	<44 (Ge $\gamma$ -assay)	<0.096
Light guides	<127 (Ge $\gamma$ -assay)	<0.015
Filler blocks	<173 (Ge $\gamma$ -assay)	<0.034
PMTs	$2.6 \times 10$	0.140
PMT mounts	7565	0.010
Rn emanation	<44	<0.081
Rn deposition (3 months)	38	0.010
Other sources		0.04
<b>Total</b>	<b>&lt;2.7x10</b>	<b>&lt;0.35</b>





# Construction Progress

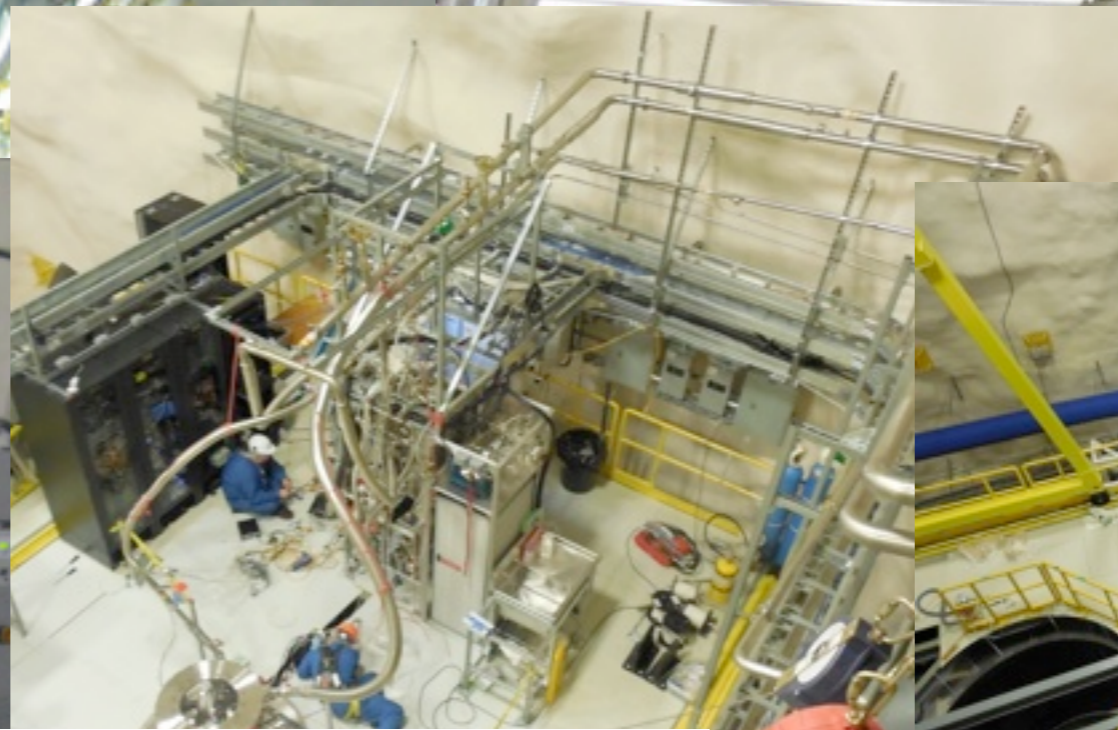
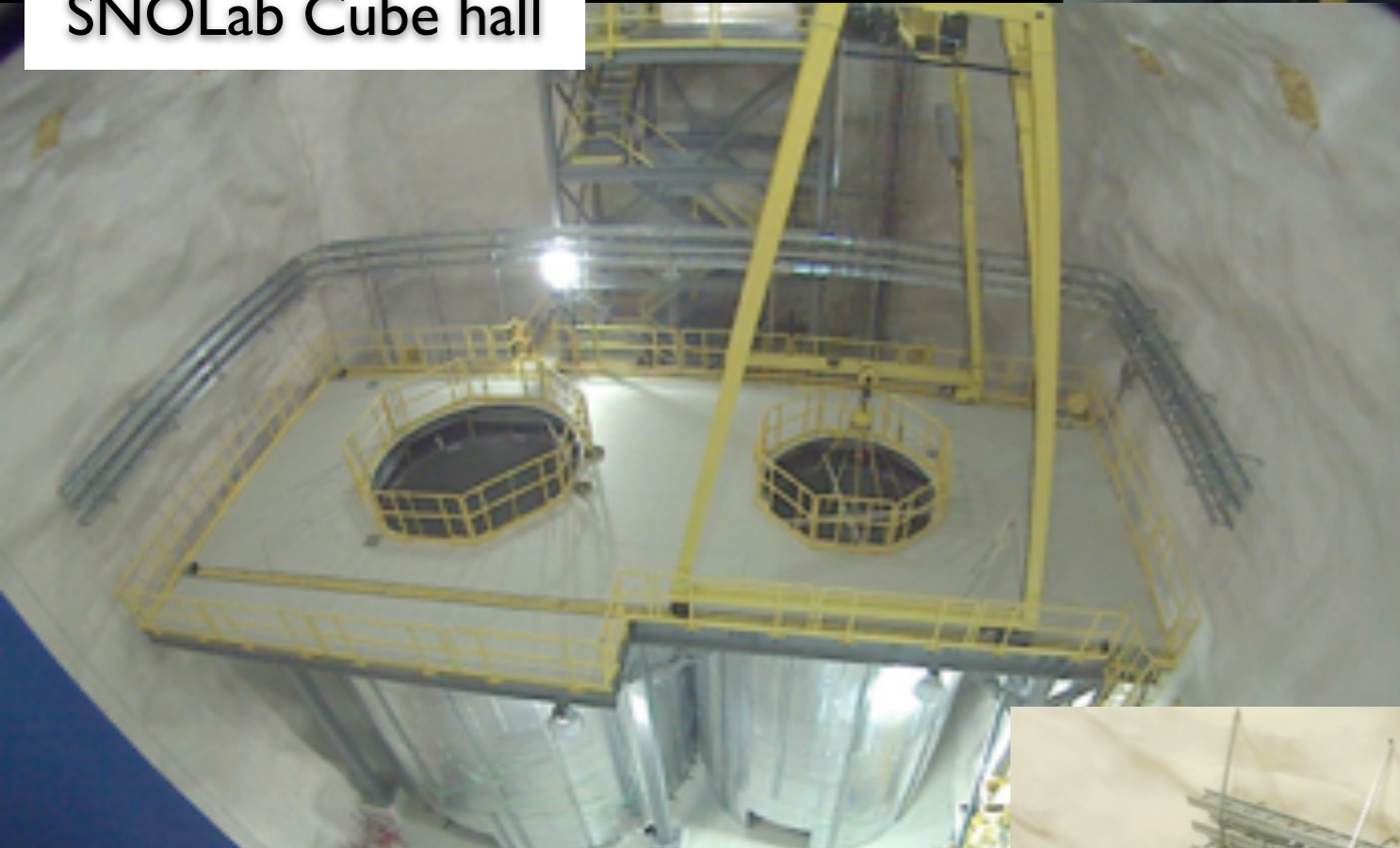
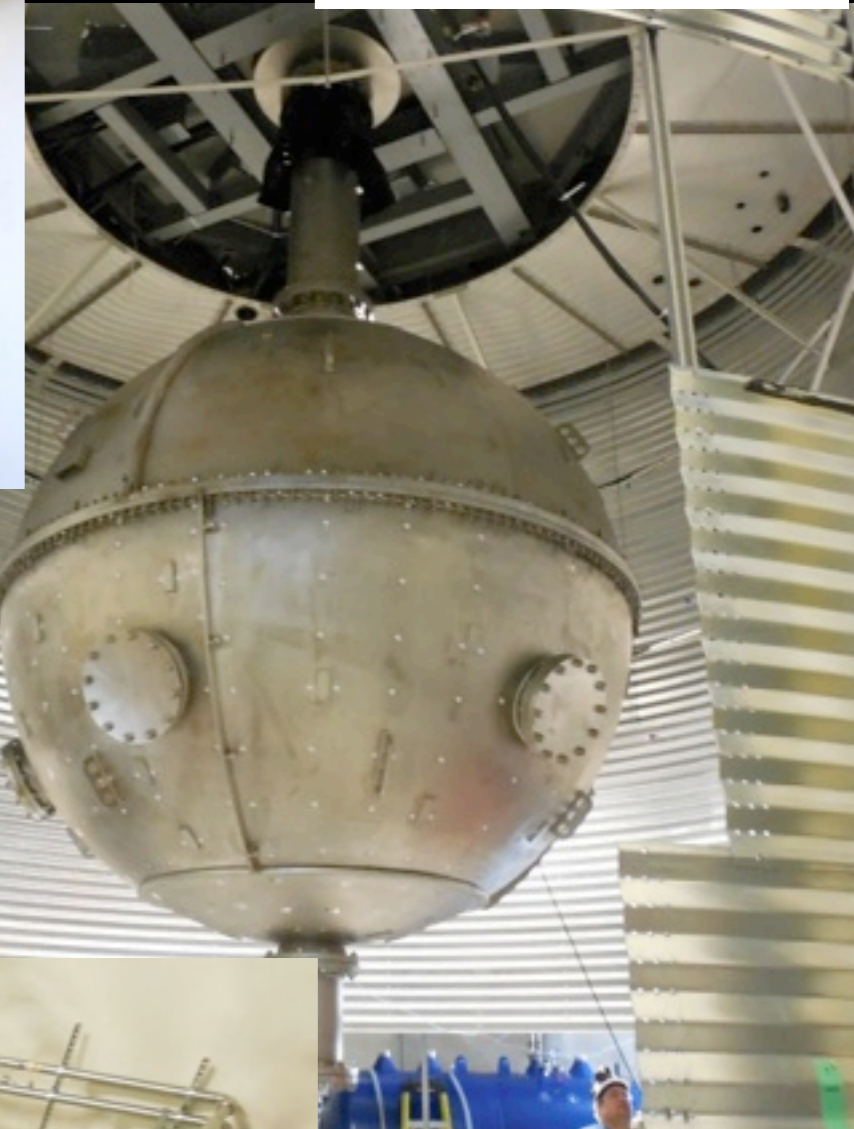
Infrastructure complete, cryo and DAQ running underground.

SNOLab Cube hall

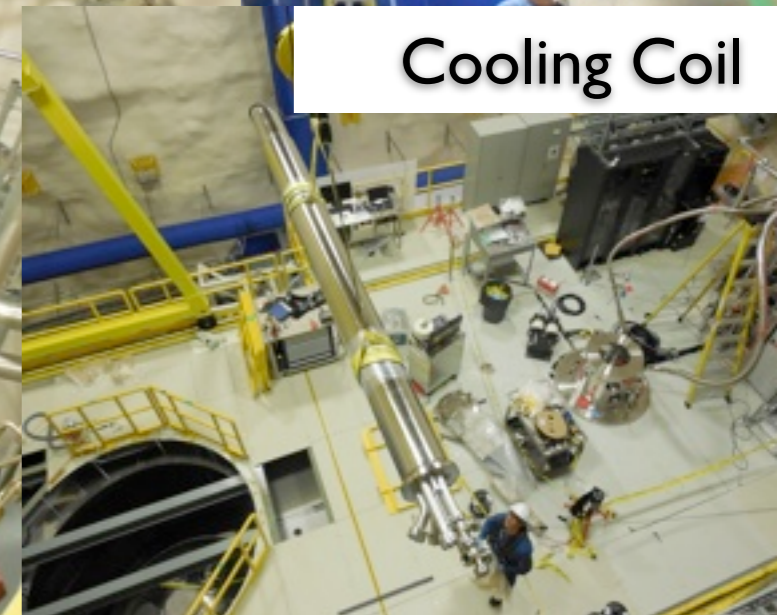
Deck Installation



Steel Shell  
in the veto tank



Cooling Coil



Process Systems and Electronics



August 22, 2014



# Construction Progress

light guide bonding

Acrylic vessel



Annealing in place



Bonding complete





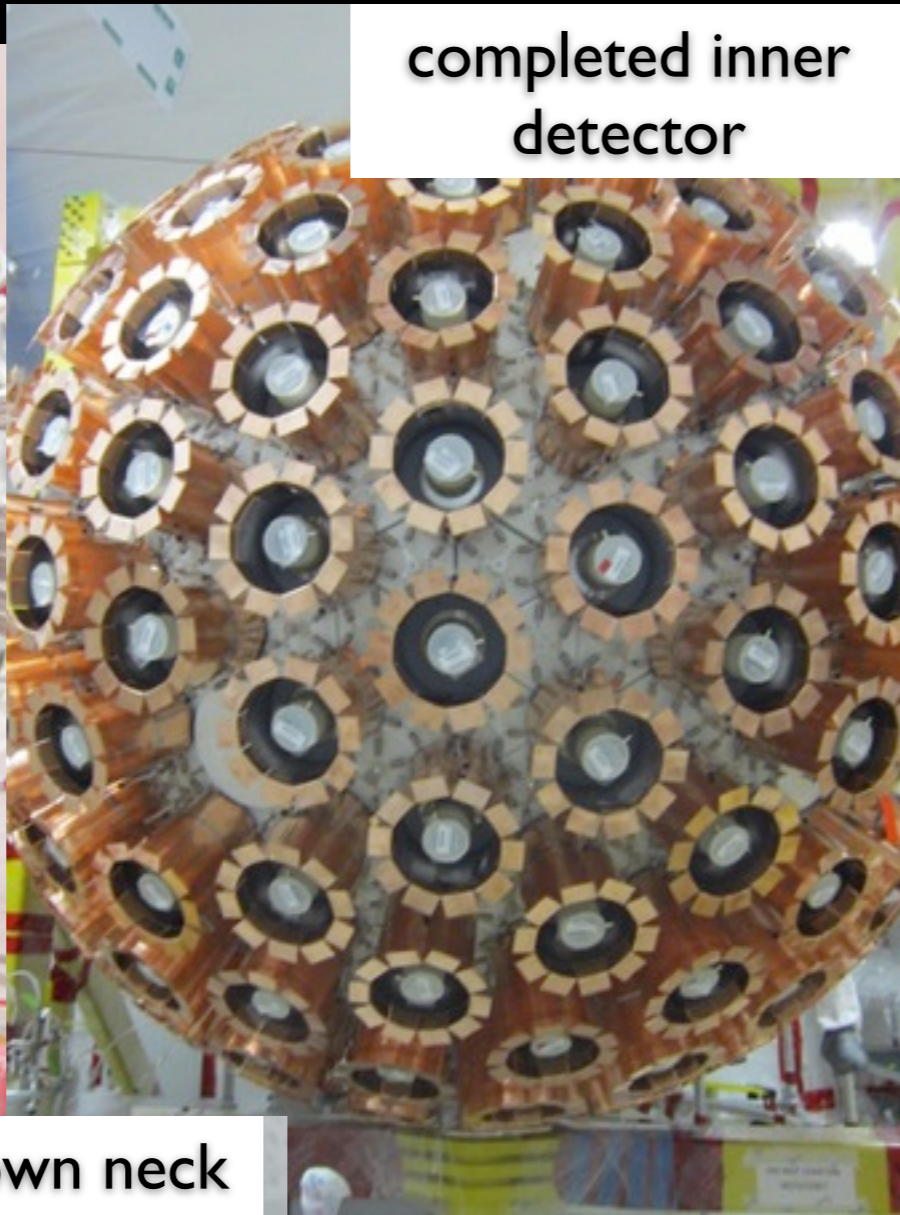
# Construction Progress

Internal resurfacing now, TPB evaporation next, LAr fill Fall 2014, physics 2015.

PMT Installation



completed inner detector



Detector Installation in Veto Tank



View down neck





# DEAP3600 Calibration System (UK)

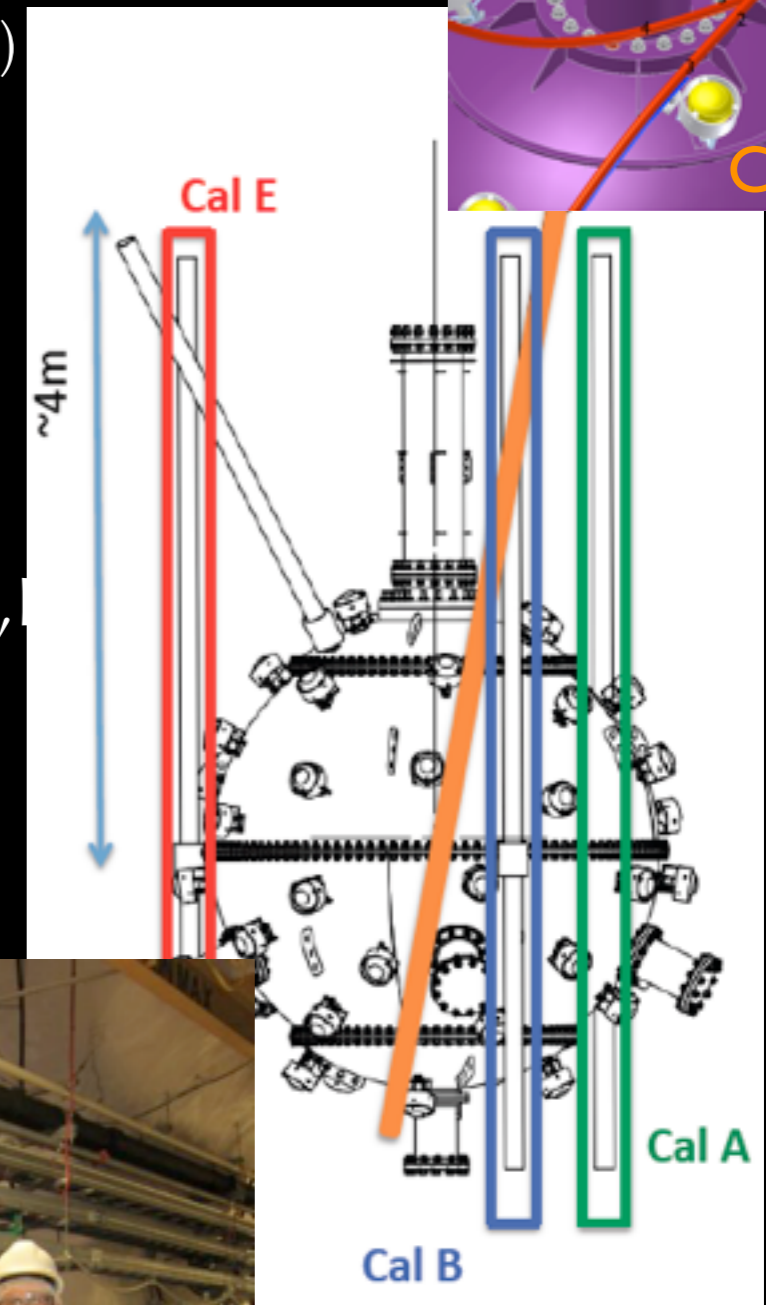
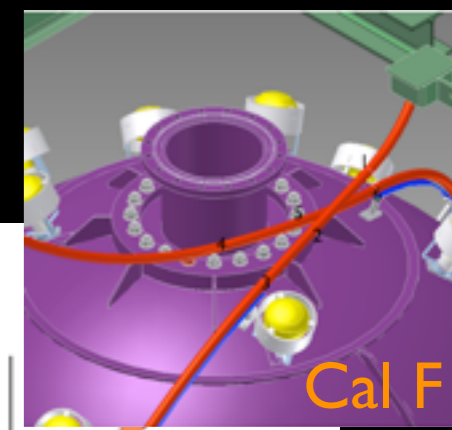
## 1. Sources and Ports:

- tagged Na-22 source Cal A,B,E pipes, Cal F racetrack (RAL)
- tagged AmBe source in Cal A,B,E (RHUL)
- optical calibration sources in neck (laser- and LED-flasks), 20 PMT lightguide reflectors (fixed), and neck laser (Sussex)

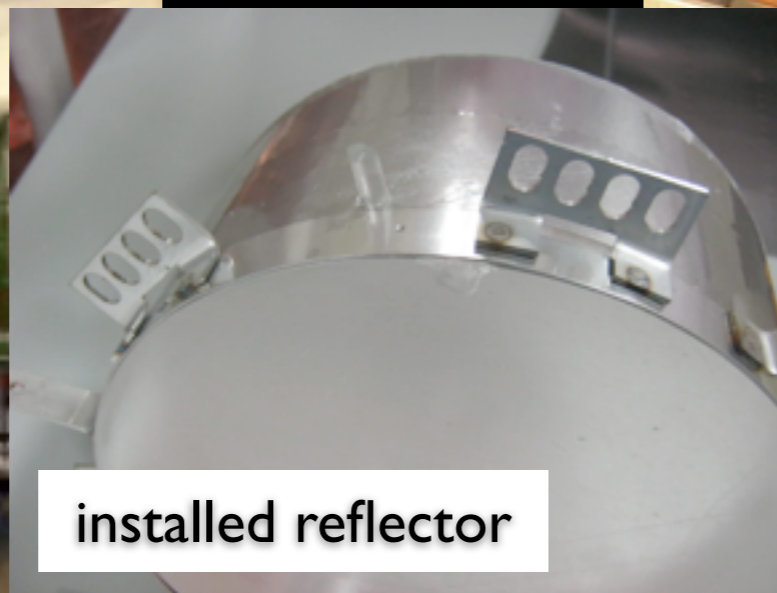
## 2. Deployment systems:

- source deployment systems for Cal F and Cal A,B,E
- neutron source deployment / HV delivery system for Cal A,B,E
- LED flask deployment through neck
- Acrylic reflector array + fibers + LED drivers

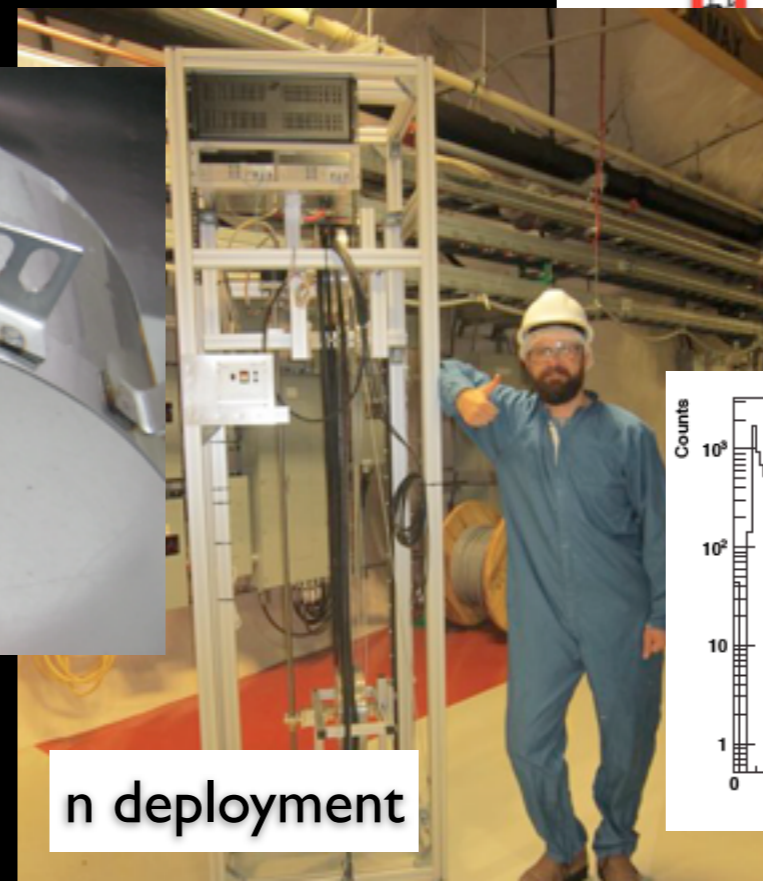
*Calibration commissioning on-site now underway!*



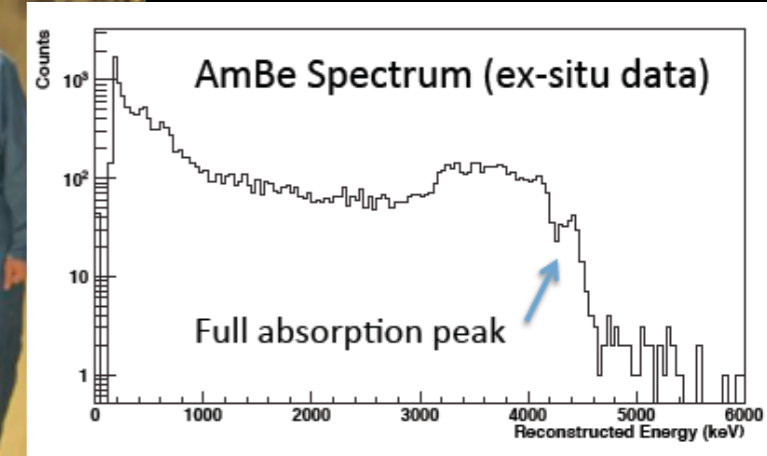
gamma system deployment test



installed reflector



n deployment



# Summary and Outlook

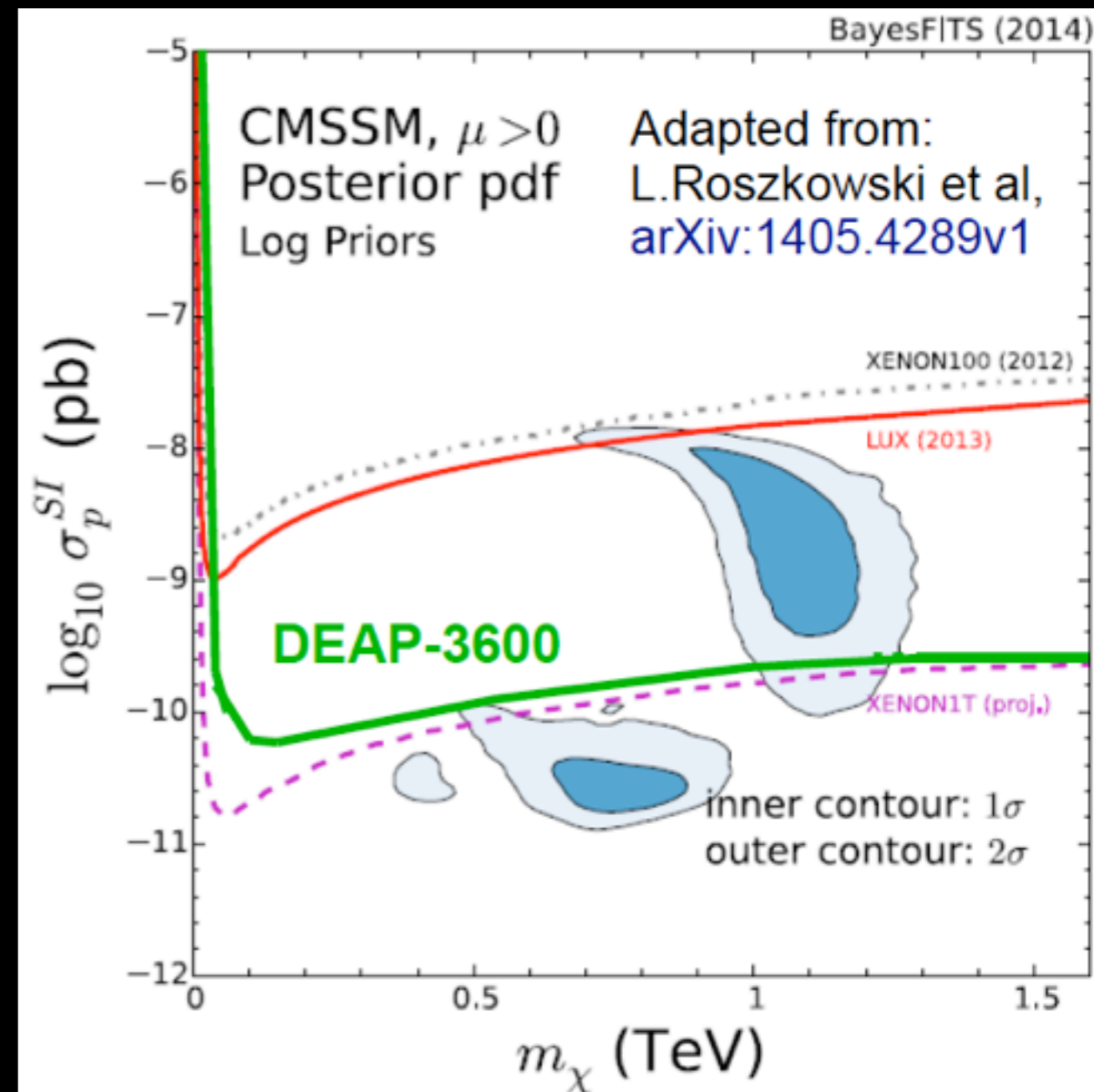
DEAP3600 will be the first demonstration of single-phase liquid Argon technology.

DEAP/CLEAN Detector development programme:

- MiniCLEAN: measure PSD, prototype LAr/LNe target exchange to test  $A^2$  scaling
- DEAP3600: dark matter discovery reach of  $10^{-46}$  cm<sup>2</sup> in 3 tonne-yrs exposure, at conservative 60 keVr threshold.
- prototypes for kT-scale, O(10s) of keV threshold detector for 'low-energy frontier' physics

DEAP3600 construction nearly complete

- resurfacing underway, TPB evaporation next, LAr fill soon!
- process systems, DAQ, calibration commissioning underway
- physics 2015!



*Stay tuned...*





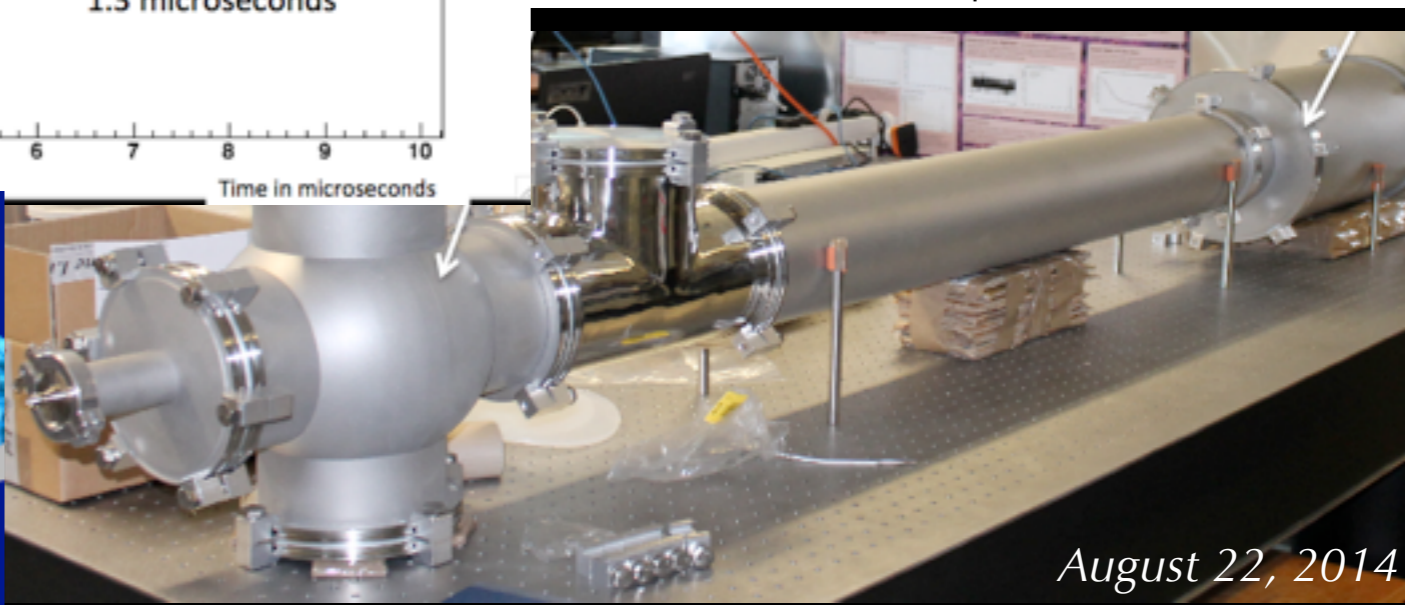
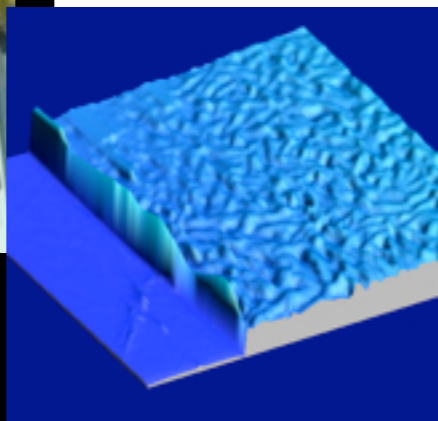
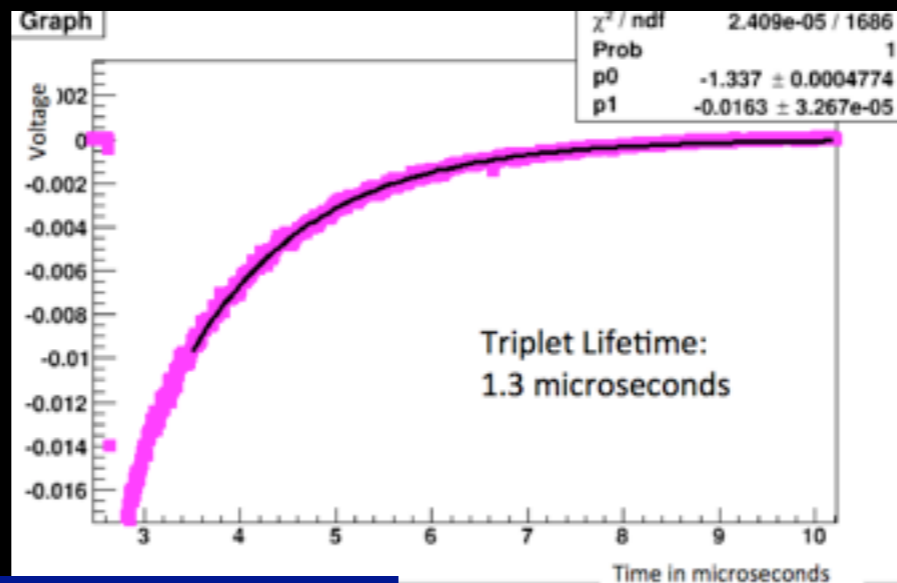
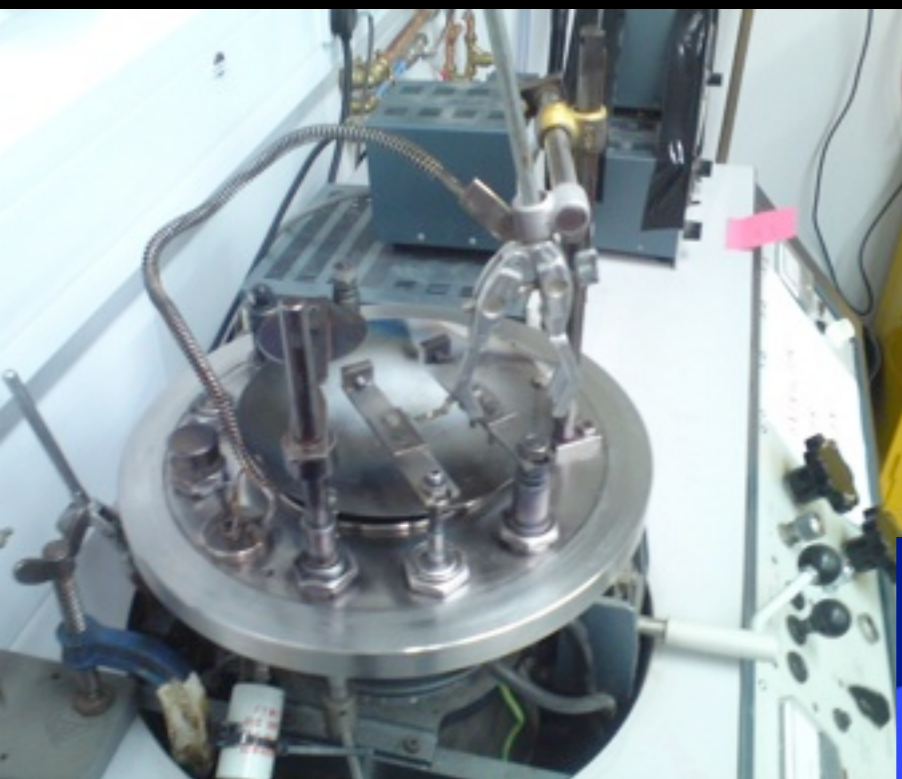
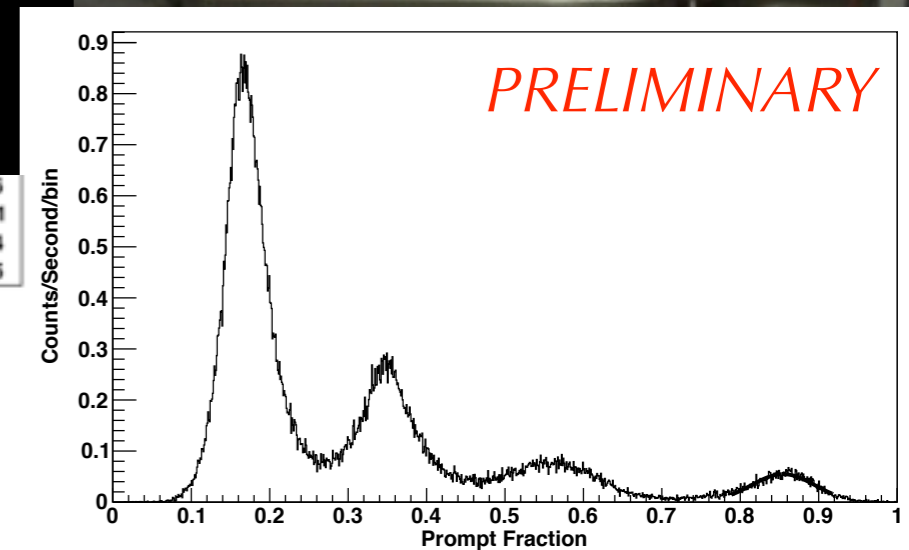
# Other Slides



# Calibration R&D (STFC PRD)

*What if we see 5 events? How will we know if this is a signal?*

- Objectives: ex-situ measurement input to calibration analysis,
  - (i) reduce systematics on energy, radius reconstruction,
  - (ii) break correlations between parameters for MC tuning
- source R&D: study modeling of source calibration (RAL)
- scintillation R&D: measure the scattering length and temp. dependence in noble liquids, explore laser calibration (RHUL)
- optical response R&D: measure the optical properties of TPB wavelength shifter (University of Sussex)





# Depleted Argon

A. Wright, *arXiv:1109.2979*

- $^{39}\text{Ar}$  beta decays with 565 keV endpoint, at  $\sim 1$  Bq/kg with half-life 269 years
- $^{39}\text{Ar}$  production supported by cosmogenic activation, underground Ar has less!
- low-background Ar sources reduce Ar-39 by a factor of  $>50$  at least (counting-only analysis)

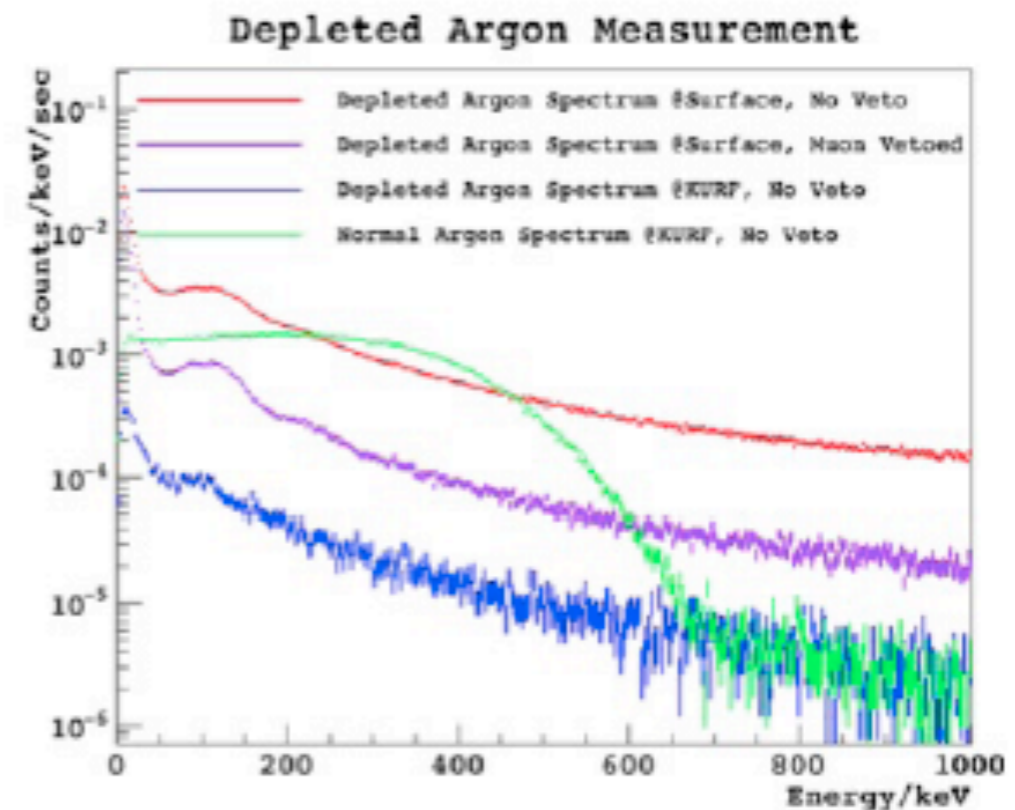
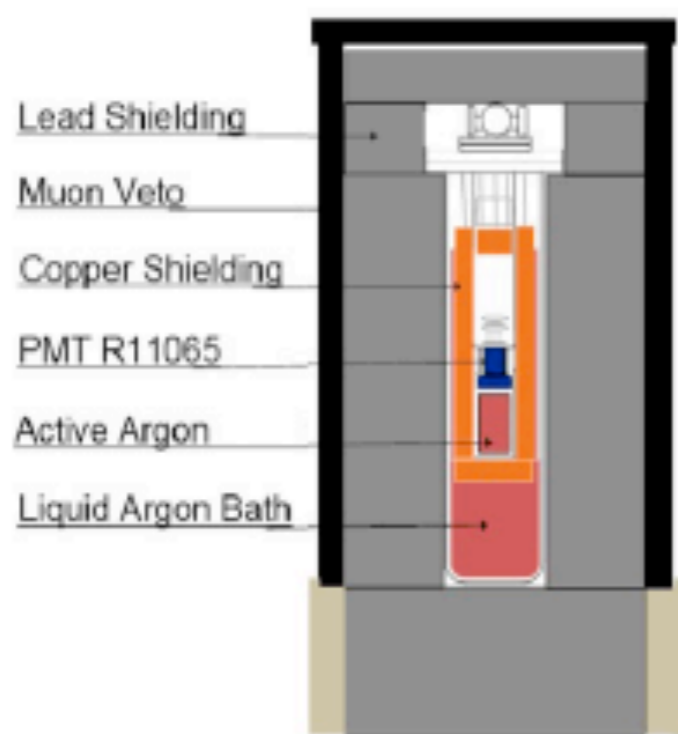


Figure 2: Left: Schematic diagram of the “Low Background Detector.” Right: The depleted argon spectra obtained in various detector configurations. In the measurement at KURF, the total event rate in 300-400 keV is  $\sim 0.002$  Hz, about 2% of the rate expected from  $^{39}\text{Ar}$  in atmospheric argon. Data taken with atmospheric argon is shown for comparison (green) - in this data the  $^{39}\text{Ar}$  spectrum is clearly visible.

