







THE UNIVERSITY of EDINBURGH

Euclid weak lensing cosmic shear measurement on IRIS/gridPP

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ESA & Ma, Ebeling, Barrett (2009)



11 Oct 2022

energy content of the universe

dark energy - responsible for the accelerated expansion of the universe - deviation from general relativity? - cosmological constant?

dark energy

69%

ordinary matter

5%

dark matter

26%

ESA



dark matter

- responsible for the rotation speed of galaxies (constant with distance from the centre)

- new particle?
- galaxies live at the centre of DM halos

95% is unknown!



size / scale



- Spacetime tells matter how to move; matter tells spacetime how to curve (J. Wheeler)
- Spacetime also tells light how to bend!







Full-sky convergence map (dark matter distribution)



Timelapse and maps made by GC using Takahashi simulations

Full-sky convergence map (dark matter distribution)



Map made by GC using Takahashi simulations

Euclid lensing on IRIS/gridPP





1/30th of the full Moon





Courtesy of L. Miller, adapted from Bridle, Kitching+ (2010)

- The shape of distant galaxies gets distorted by the large scale structure (cosmic shear)
- Our goal is to recover the cosmic shear through the observation of blurred, pixellated images
- Challenging task: do it for 1.5 billion galaxies (my main job in the project)
- LensMC official measurement method





Data sent to ground for analysis

Expected ~10PB of data (~10,000 laptop storage)

Data analysed by centres spread across Europe

 Institute for Astronomy: leading the scientific analysis and exploitation of such sheer volume of data

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Challenges

- Cosmological measurements with percent-level precision: measure a cosmic shear of 1% with an accuracy of 0.2%
- Correct for PSF errors ~10⁻⁴
- Galaxies down to S/N>5, 30/sq.arcmin, median z=2
- Need 10⁸-10⁹ galaxies to beat down shape noise (~0.3), g=<e>
- Euclid:
 - 15,000 sq.deg across the extra galactic sky (37% of the full sky)
 - Mag<24.5 for wide survey; +2 for deep survey
 - Diffraction limited with a resolution of 0.1"
- Estimate 2p-correlation/ power spectra of:
 - Weak Lensing: shape-shape
 - Galaxy Clustering: position-position
 - Cross-correlation: shape-position

Science objectives



How can IRIS and GridPP help?

- Lots of help through many cores up and down the country
- Lots of CPUh to reduce measurement uncertainty
- Demonstrate performance of LensMC shear measurement method (the official method) at Euclid requirement level
 - very hard to meet with clusters
 - in fact, been using IRIS resources (e.g. Cambdrige, RAL) for testing, but now need to scale up
 - cores made available by IRIS to gridPP will be crucial

LensMC-gridPP

- Take large simulation of the sky
- Split in small patches of ~9 sq.arcmin
- Emulate Euclid images and run measurement
- Compile all results
- Estimate bias/errors
- Meet requirements

LensMC-gridPP

- The job:
 - take in small patch of sky, simulate image, and run measurement
 - Conda environment installer, code, Python script, wrapper script, and ancillary data on storage
 - jobs retrieve data before running, results saved back to storage
- Successful at running all of this for a couple of months on gridPP VO (70% complete):
 - very low memory footprint (~1GB), runtime ranging from few hours to few days
 - often 10,000 cores/day with very low downtime or failure rate
- Moving on:
 - new Euclid VO for tracking/accountability
 - Any further help from IRIS very much appreciated!
- Outcome: key Euclid publication (with gridPP+IRIS stamp) ready for submission to the board hopefully by the end of the year

Thank you

coming up in 2023...

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